

A Time-varying Radiometric Bias Correction for the TRMM Microwave Imager

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Thesis defense : Oct 29, 2008



Outline

- Dissertation Objective
- Initial research
 - Introduction to GPM and ICWG
 - Inter-satellite radiometric calibration
- Current research
 - TMI time-varying bias correction
- Conclusions and Future work

Dissertation Objective

- Original Objective
 - To develop inter-calibration techniques for a constellation of satellite radiometers over ocean, Amazon and polar regions
- Current Work
 - Detection of systematic errors using inter-calibration and development of correction techniques that eliminate systematic errors detected in TMI

Motivation

- GPM contains a constellation of similar, but not identical, radiometers
 - Satellites of opportunity
- Inter-calibration is required in order to obtain self-consistent retrievals from all the radiometers in the constellation
- Separating long term environmental change from instrument errors due to aging

Calibration for GPM constellation

- GPM observatory : Non sun-synchronous tropical orbit
- Constellation satellites : Sun-synchronous polar orbits



ICWG Objectives

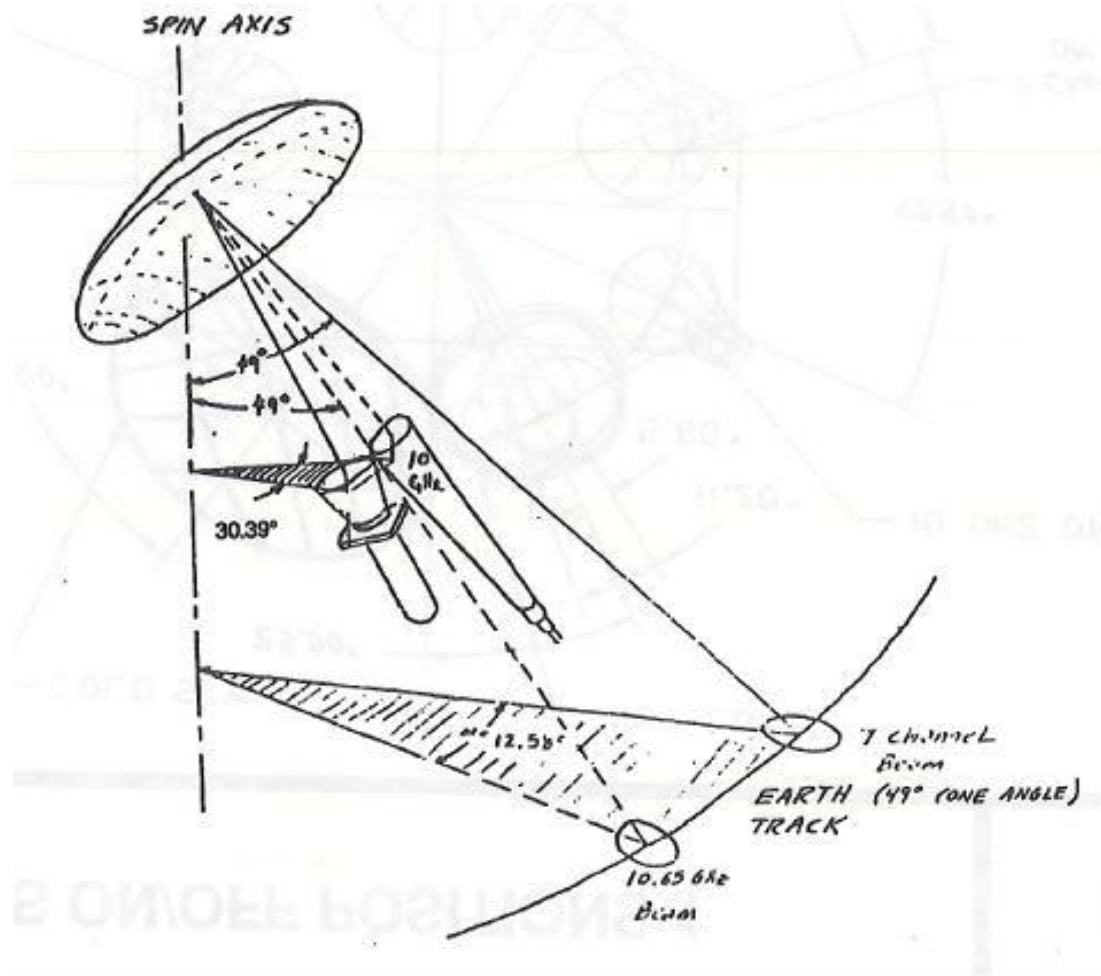
- Common Dataset
 - TMI as proxy for GMI
 - WindSat and SSMI on F13, F14 are the polar orbiting radiometers
- Verify the accuracy of TMI as the transfer standard
- Using the “gold standard” radiometer to calibrate polar orbiting radiometers
- Generation of a standard, repeatable protocol for use in GPM

ICWG Methodology

- Four teams using independent techniques
 - BESS : Tom Wilheit
 - CSU : Chris Kummerow and Wes Berg
 - Michigan : Chris Ruf
 - UCF : Linwood Jones
- Common radiative transfer model
- Common dataset : July 2005 – June 2006
- Contributions by Fuzhong Weng and others...



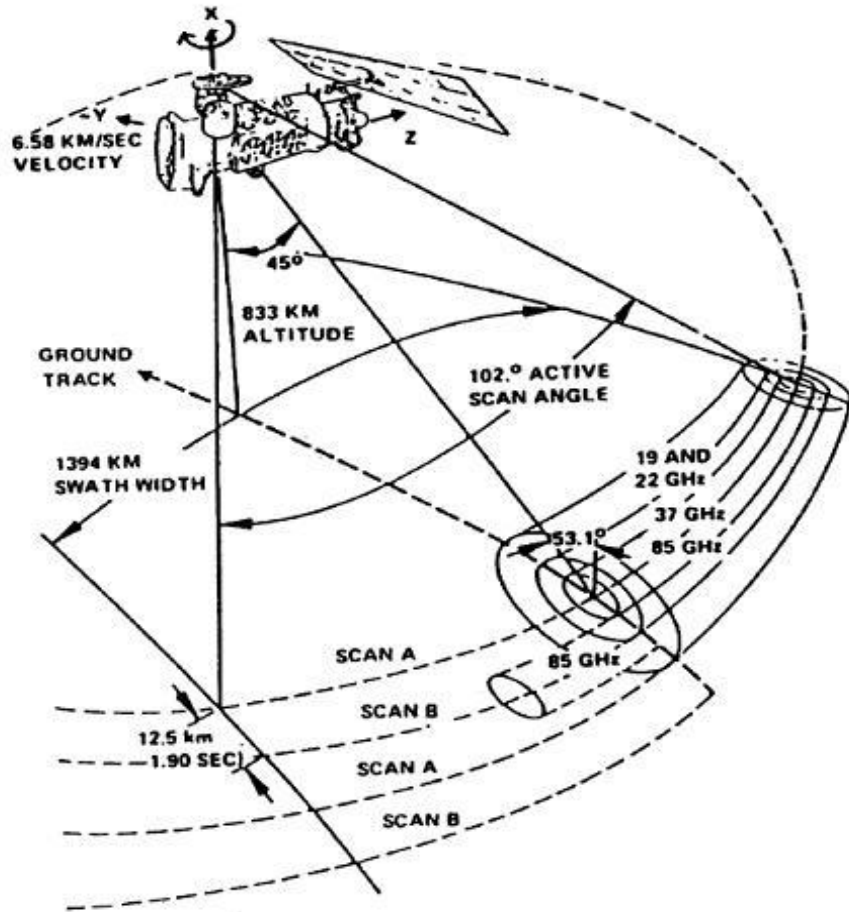
TMI Scan geometry



TMI channels

Freq GHz	Channels	EIA, deg	BW, MHz
10.65	V,H	52.3	100
19.35	V,H	52.3	200
21.3	V	52.3	500
37.0	V,H	52.3	2000
85.5	V,H	52.3	3000

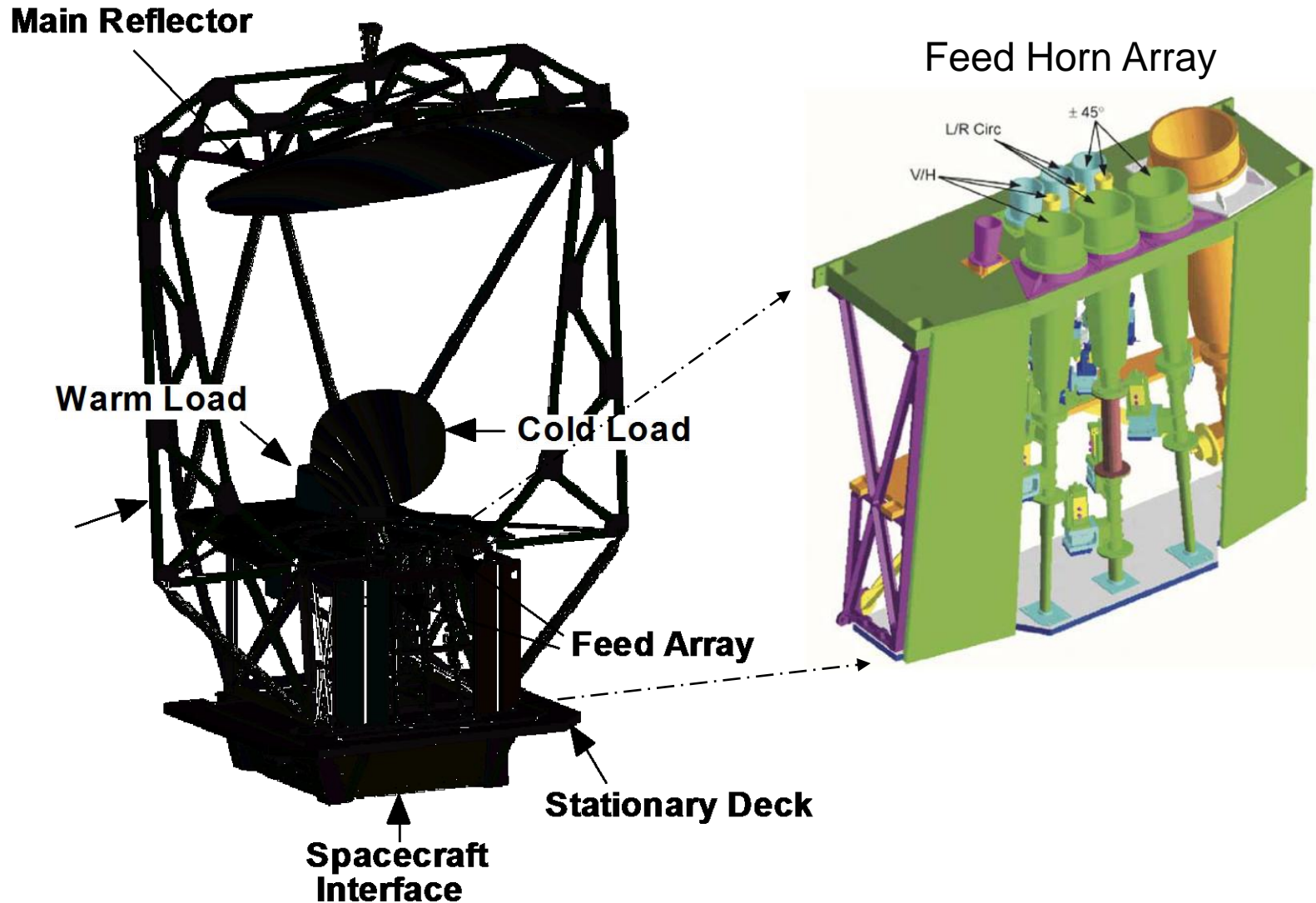
SSM/I scan geometry



SSMI channels

Freq GHz	Channels	EIA, deg	BW, MHz
19.35	V,H	53.4	250
22.3	V	53.4	250
37.0	V,H	53.4	1000
85.5	V,H	53.4	1500

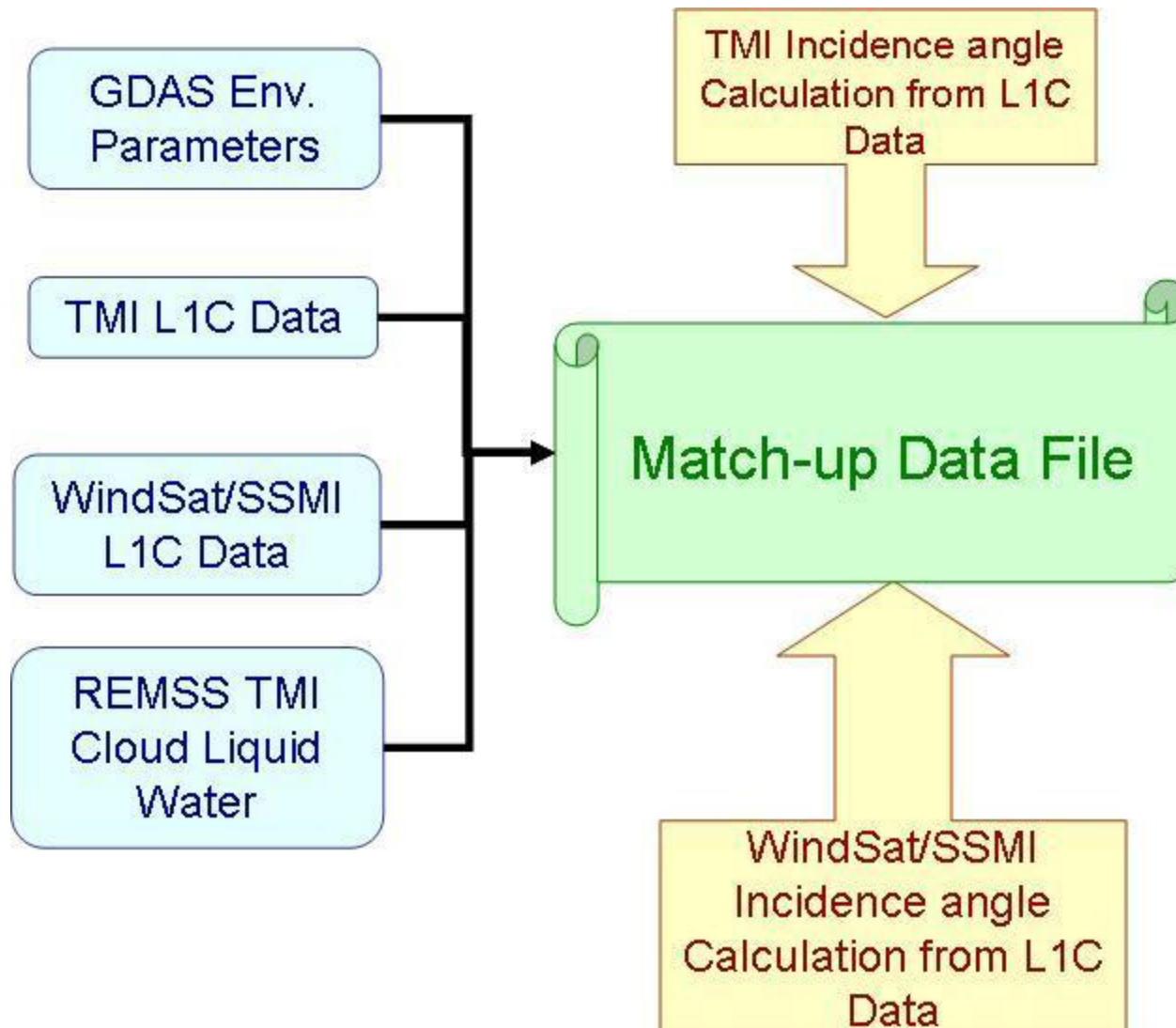
WindSat Sensor Assembly



WindSat channels

Freq GHz	Channels	EIA, deg	BW, MHz
6.8	V,H	53.5	125
10.7	V,H, \pm 45,lc,rc	49.9	300
18.7	V,H, \pm 45,lc,rc	55.3	750
23.8	V,H	53.0	500
37.0	V,H, \pm 45,lc,rc	53.0	2000

Match-up Data Sets



Match-up Data Sets

- ASCII files containing Tb meas, incidence and azimuth angles, & enviro parameters
- Match-ups with temporal tolerance of ± 1 hour and spatial quantization of 1° Lat x 1° Lng
- Three collocation files generated per day
 - TMI-WSat, TMI-SSMI-F13 & TMI-SSMI-F-14
- Match-up file formats available via SSH to the research community

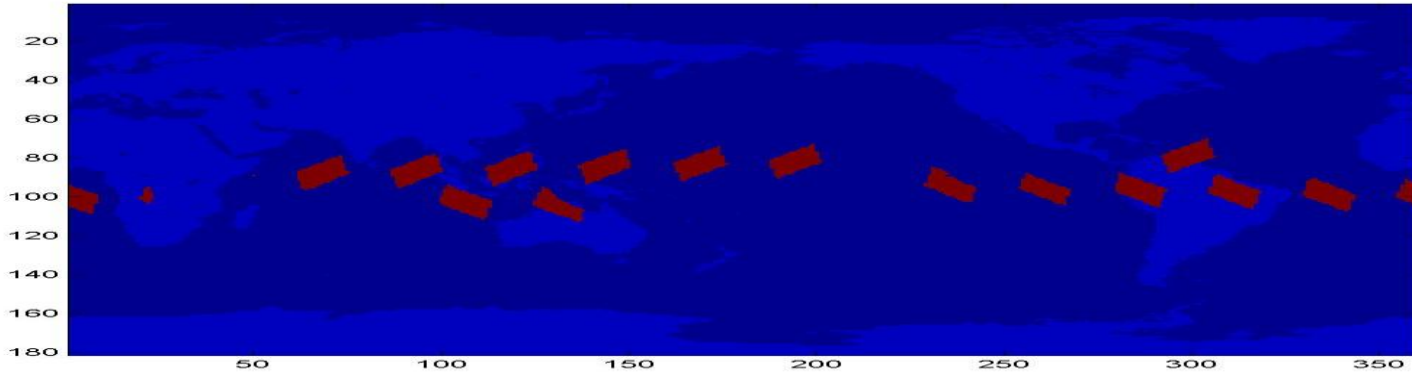
Match-up Environmental Data Sets

- Atmospheric parameter profiles from GDAS
 - Pressure, temperature, water vapor
 - @ 21 atmos pressure layers
 - Global @ 1° lat/lon grid, every 6 hours
- Oceanic surface parameters from GDAS
 - Sea surface temperature, wind speed & direction
- Ocean salinity from NOAA climatology
- Columnar cloud liquid water from TMI retrievals
 - Remote Sensing Systems

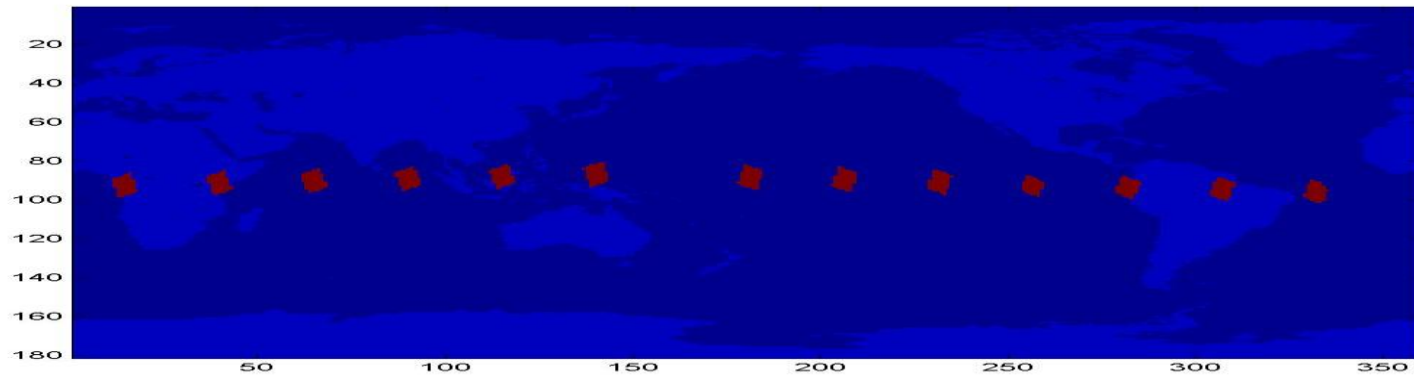


Match-up locations

SSMI F-13



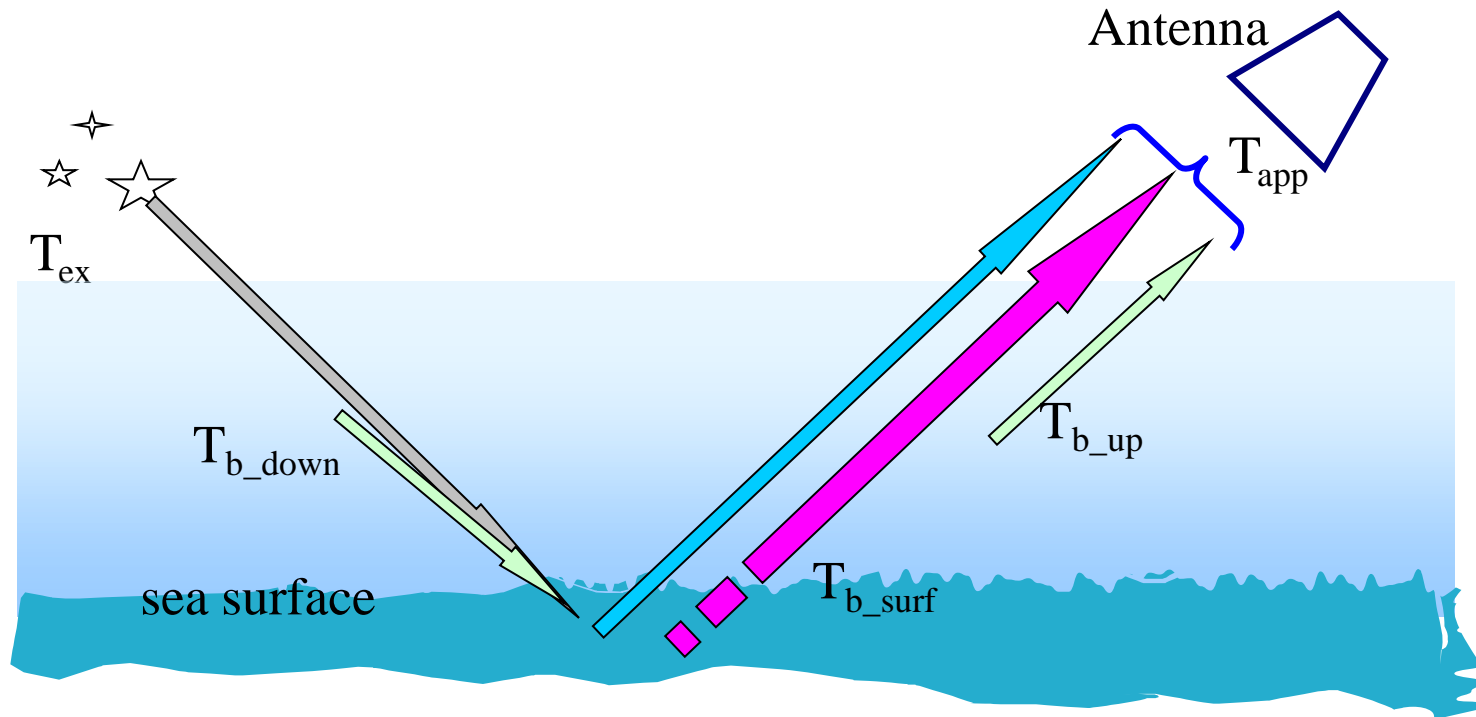
WindSat



Match-up Data Sets : Tb Filtering

- Filtering of match-up **Tb** samples to remove non-homogeneous environmental scenes & bogus outliers
 - In each 1° lat/lon bin, std dev of:
 - V-pol $\leq 2K$, H-pol $\leq 3K$ Upper limits of Tb's applied to remove rainy and land pixels
 - TMI: 10V (185K), 10H (115K), 19V (230K), 19H (200K), 21V (260K), 37V (240K), 37H (210K)
 - WindSat: 6V (200K), 6H (120K), 10V (200K), 10H(150K), 18V (250K), 18H (200K), 23V (260K), 23H (230K), 37V (250K), 37H (200K)
 - Conservative “Land Mask” applied

Radiative Transfer Theory



$$T_{refl} = (1 - \varepsilon)(\tau T_{ex} + T_{b_down})$$

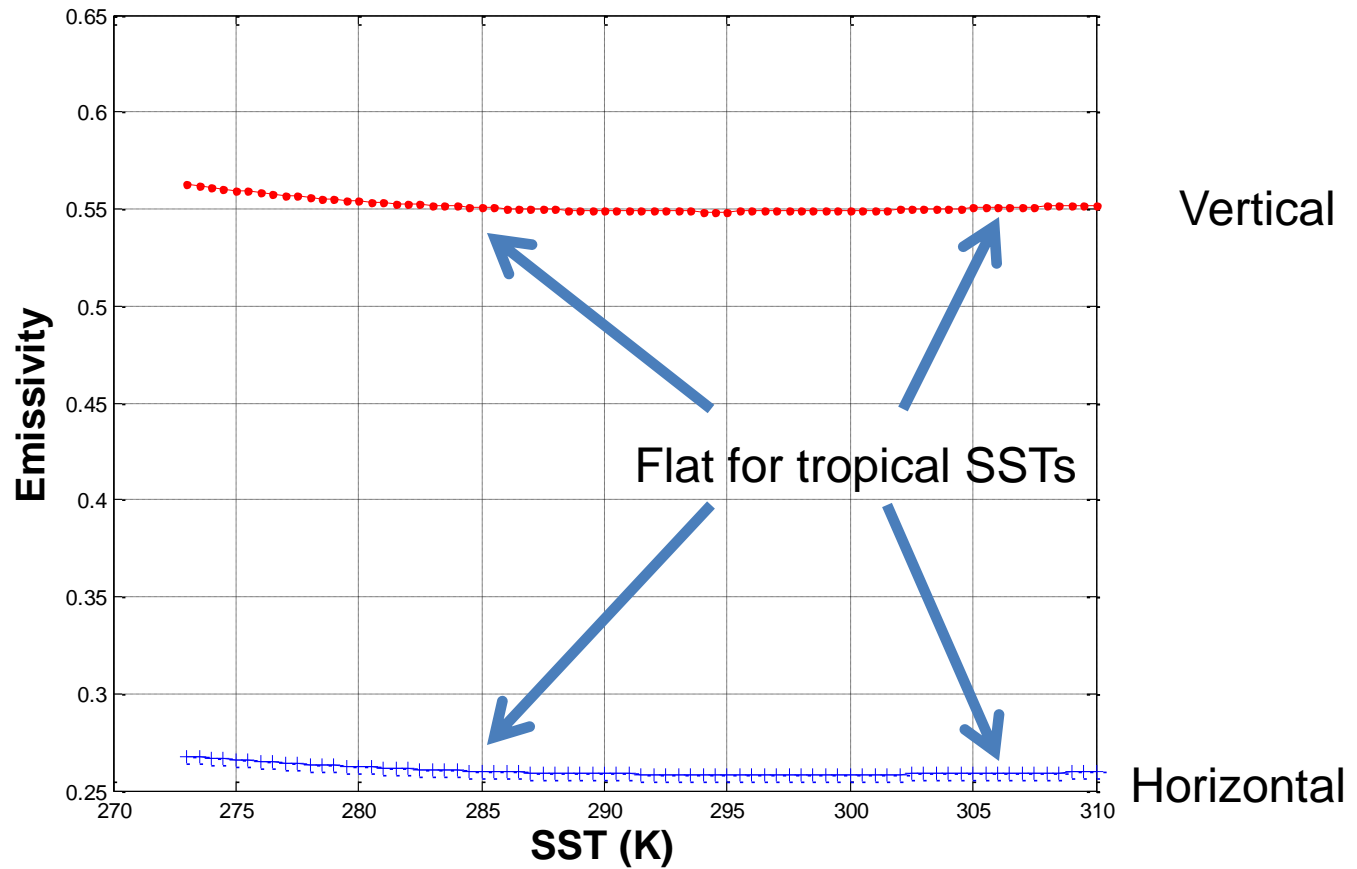
$$T_{app} = T_{b_up} + \tau(T_{b_surf} + T_{refl})$$

Radiative Transfer Model - ICWG Subroutines

- 100 atmos layers, each of thickness 200m
- Elsaesser model for ocean isotropic emissivity
- Weng model for ocean directional emissivity
 - Anisotropic with relative direction
 - Calc from GDAS wind dir & Tb observation azimuth
- Rosenkranz model for WV, CLW, O2 and N2 absorption
- Implemented in MATLAB and C



SST dependence of 10.7 GHz emissivity



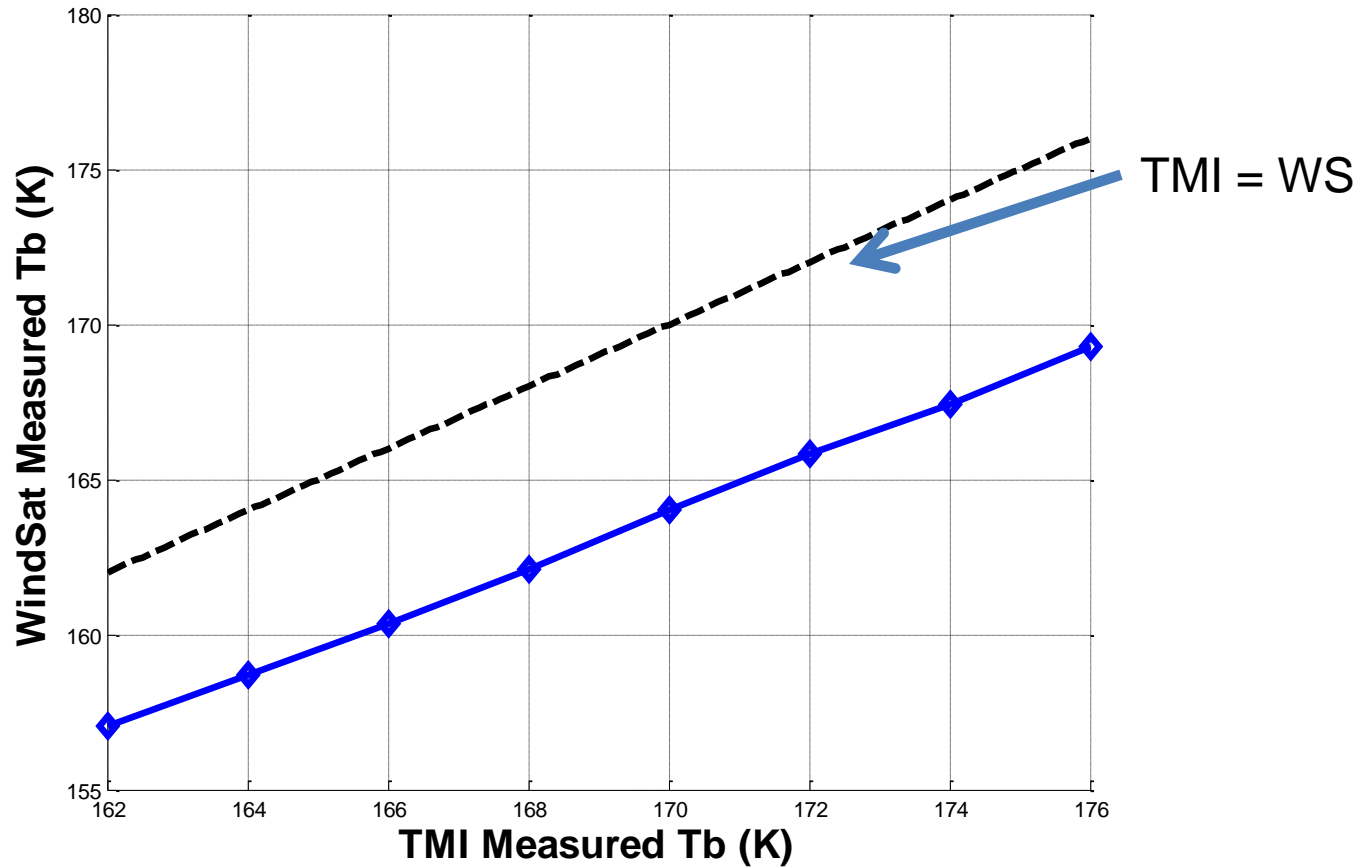
TMI/WindSat Radiometric Biases

$$Tb_SS_{norm} = Tb_SS_{obs} + Tb_TMI_{pred} - Tb_SS_{pred}$$

$$Tb_diff = Tb_TMI_{obs} - Tb_SS_{norm}$$

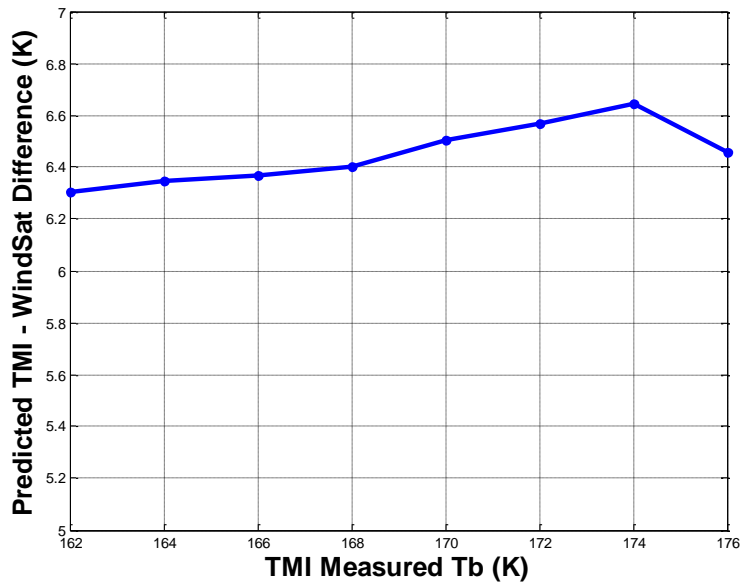
- Tb_SSnorm : WindSat/SSMI Tb normalized to TMI freq. and angle
- Tb_diff : Unexplained radiometric bias between TMI and WindSat/SSMI

Example : 10V

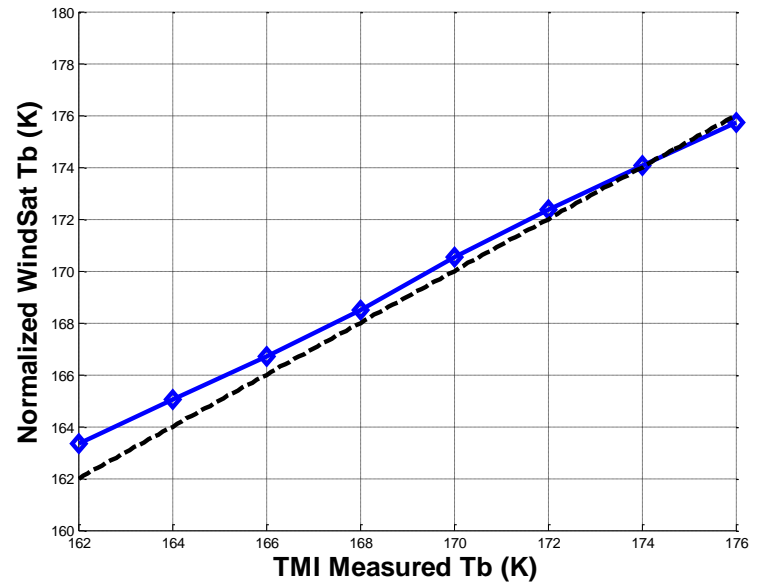


Example : 10V

Predicted Tb difference



Normalized Tb

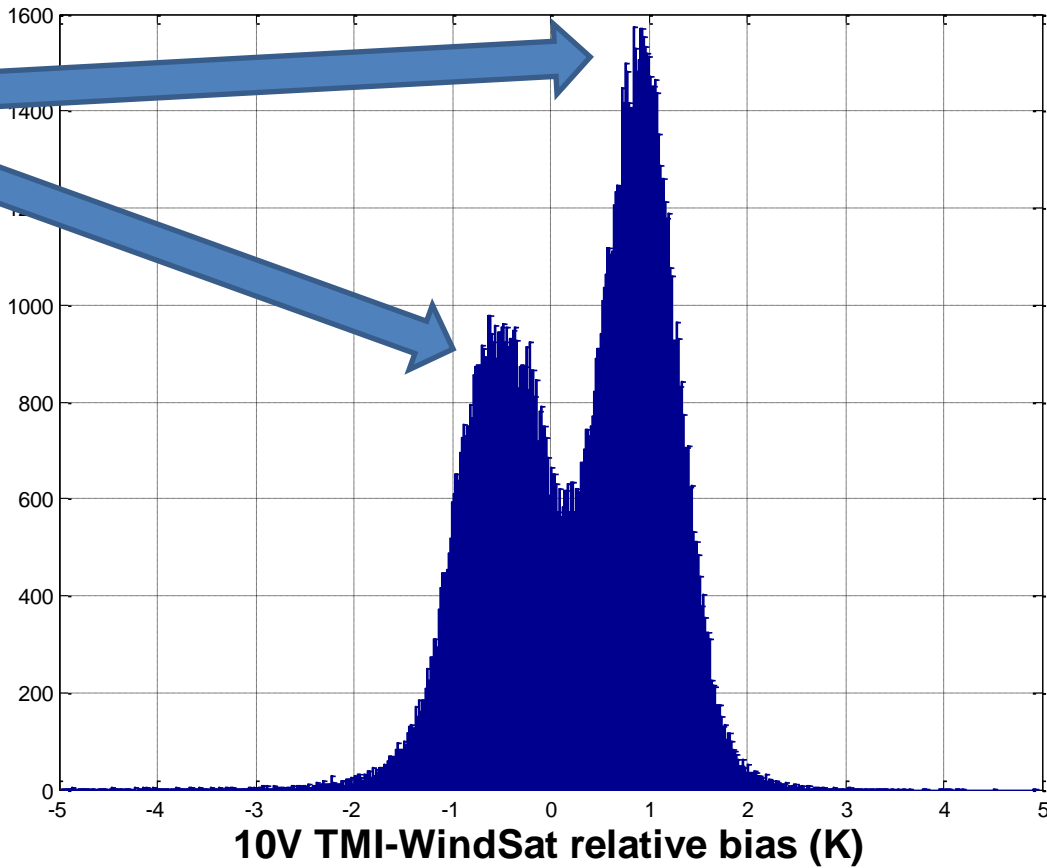


Inter-Sat Tb Comparison Results

- Relative Biases for TMI and WindSat analyzed for 1 year (ICWG period)
- The radiometric bias between TMI and WSat radiometers **must be** independent of enviro parameters
- Anomaly:
 - Results indicate apparent correlation between the bias and SST.

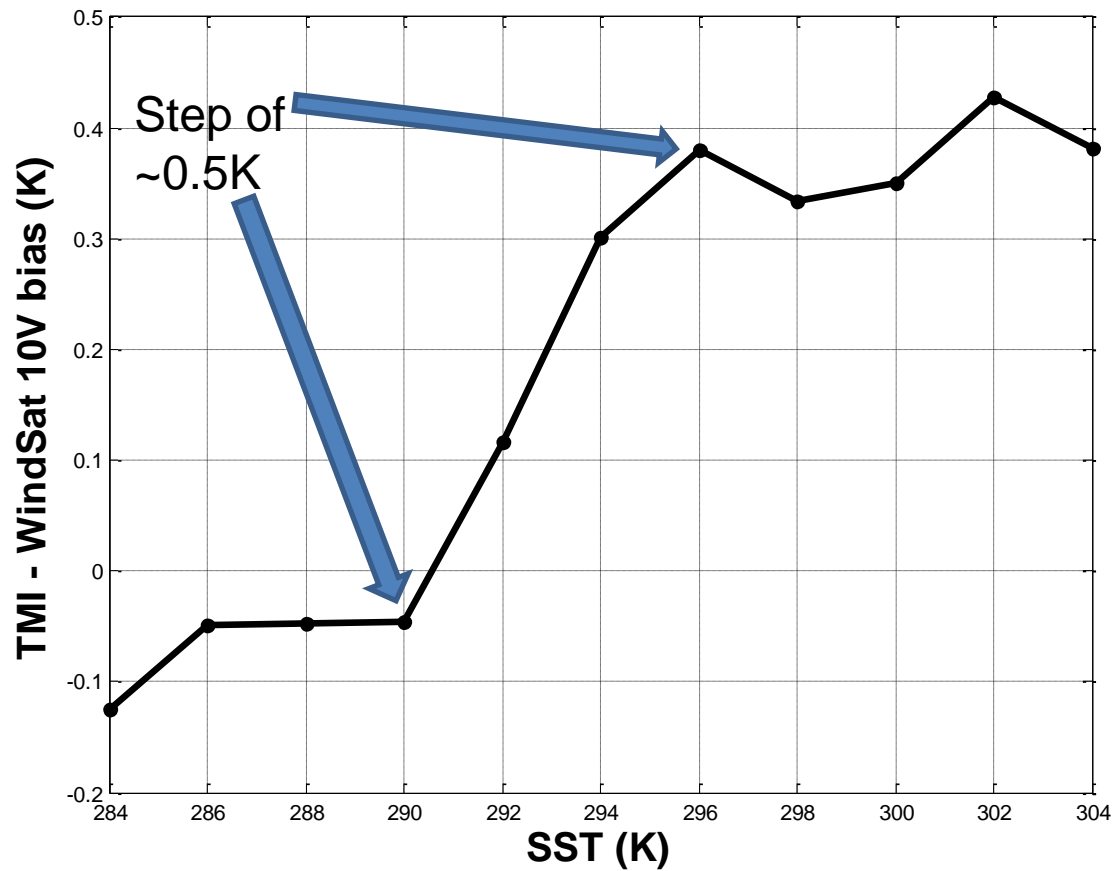
Inter-Sat Tb Comparison Results

Two distinct
random error
sources?

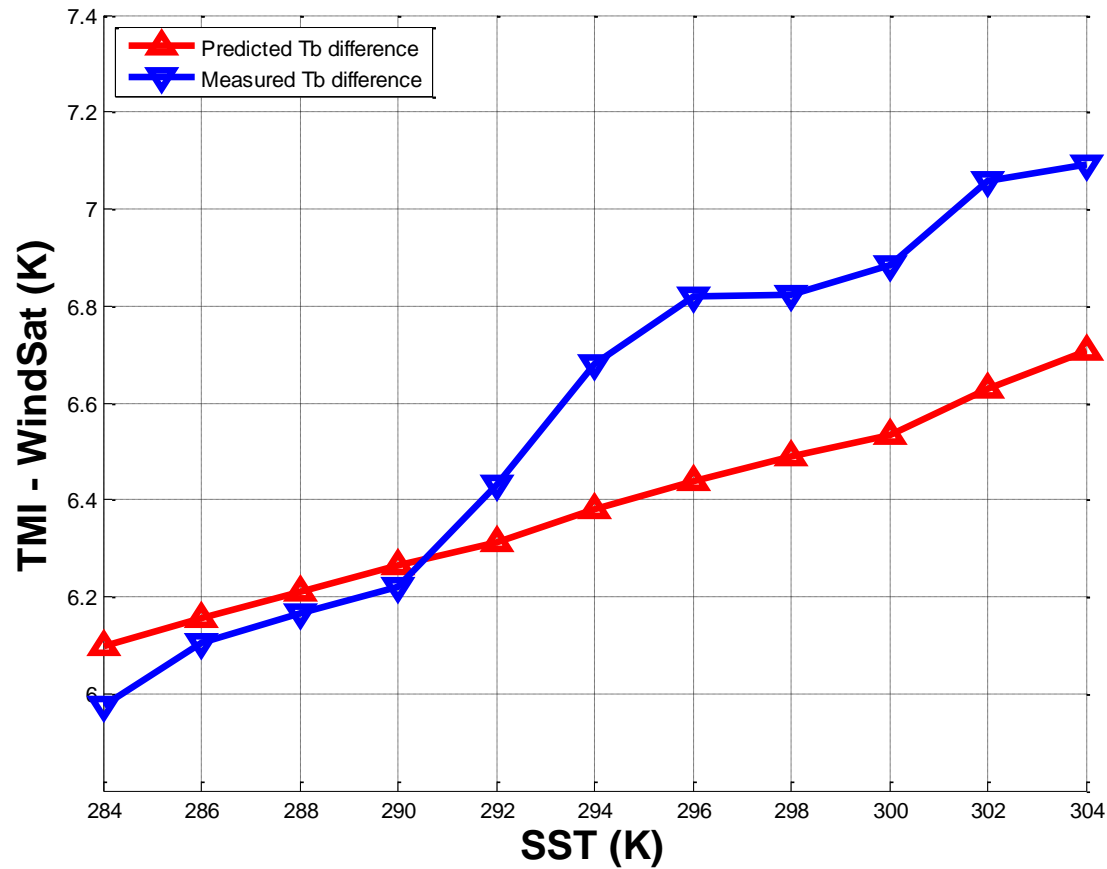


10V Bias between WindSat & TMI

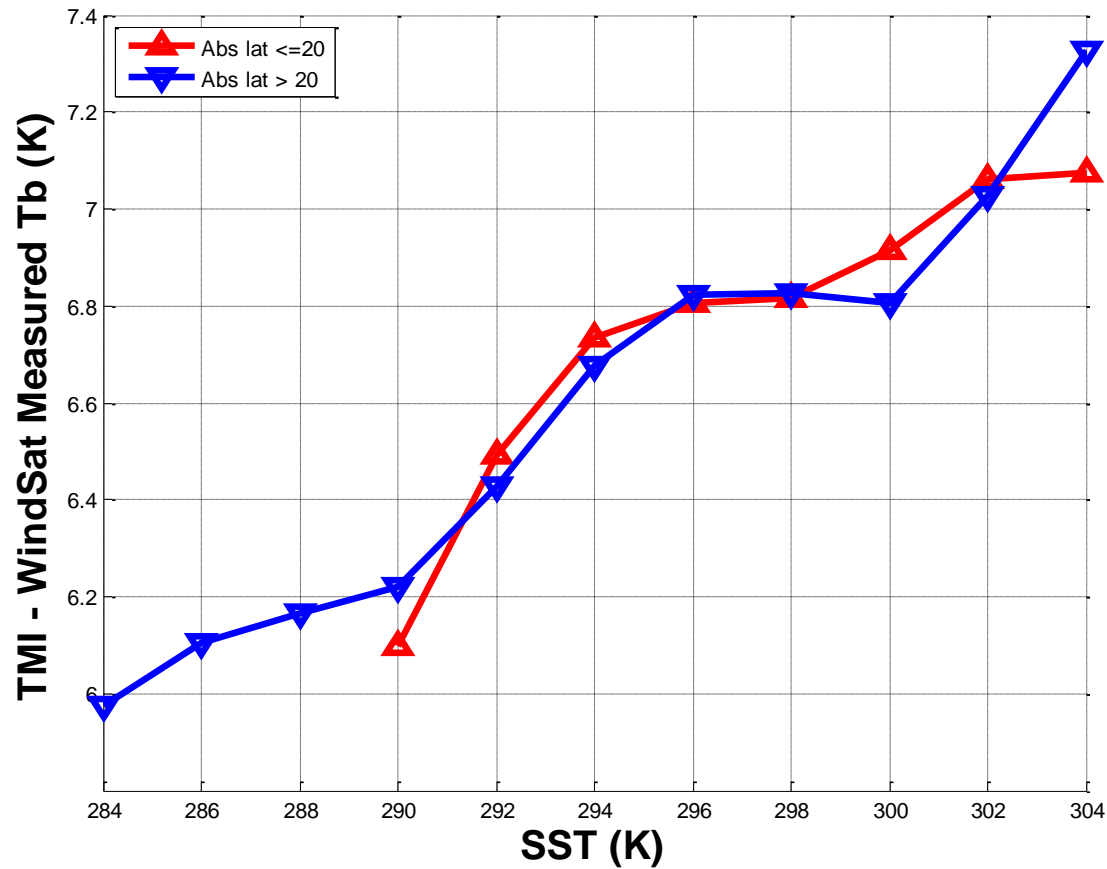
Concern: bias is $f(\text{envir pars})$???



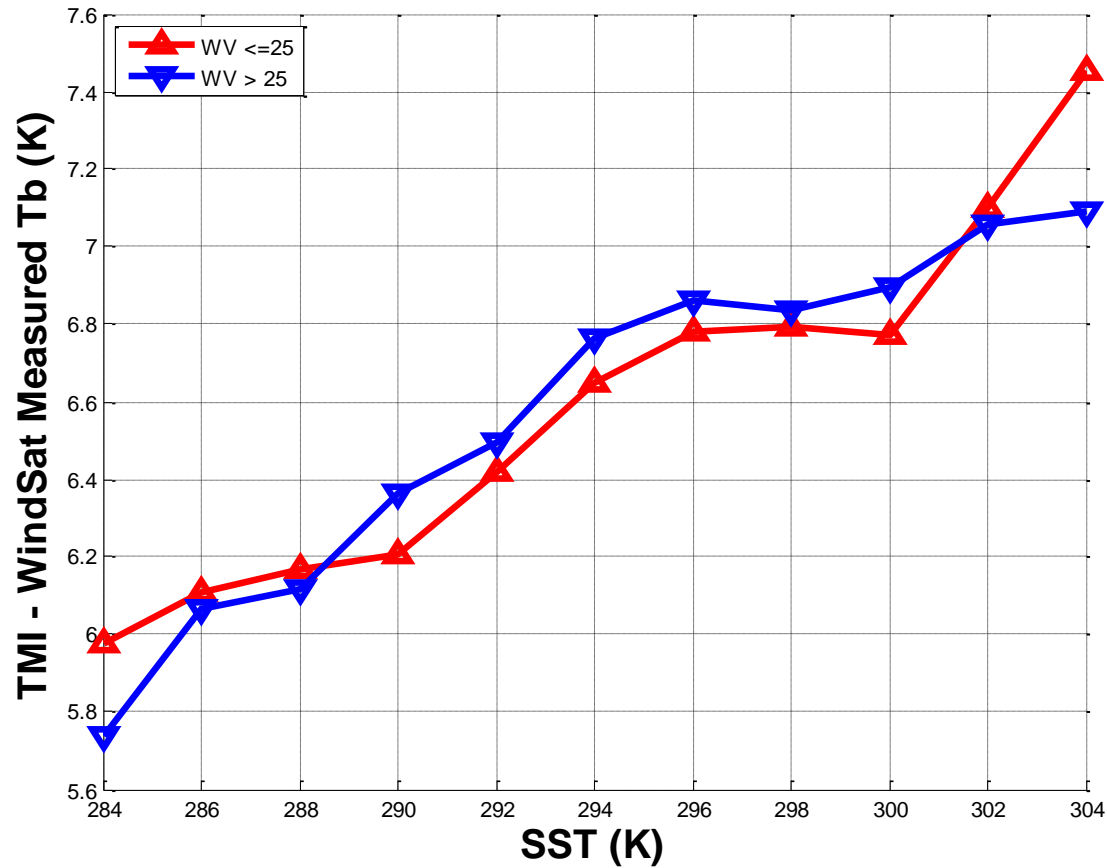
Predicted and measured differences



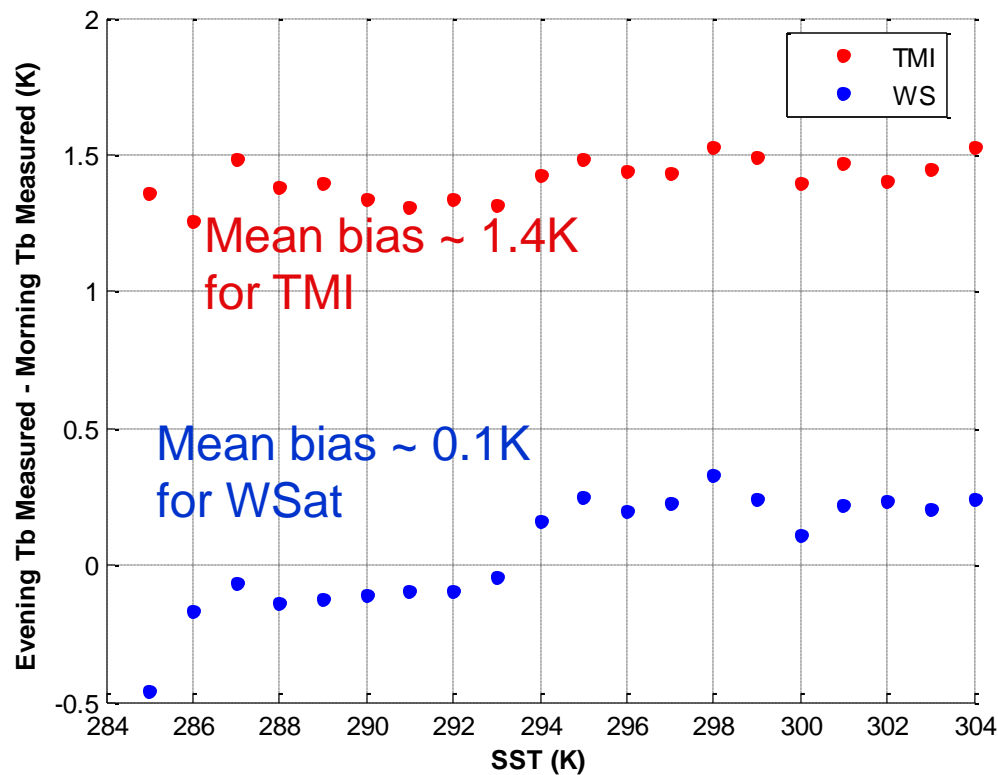
Latitude dependence of measured Tb differences



WV dependence of measured Tb differences



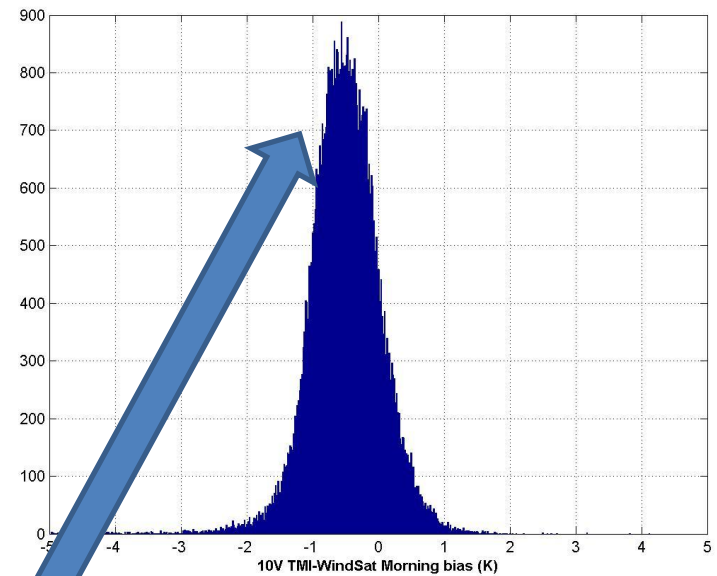
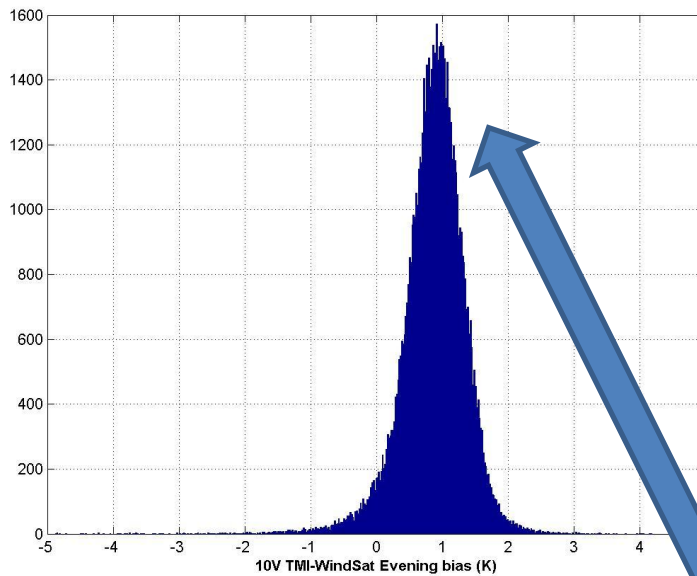
(Evening – Morning) measured Tb diff.



Evening/Morning separation

Evening

Morning



Unimodal

TMI Time of Day bias variation

- TMI reflector has been estimated to have a reflectivity of about 96% [Wentz et. al. 2001]
 - Bias varies with reflector physical temperature, which has a diurnal cycle
- Challenge : WindSat, F13 and F14 have very similar equator crossing times
 - TMI daily bias variation cannot be estimated from the match-up dataset
- Solution : Compare TMI meas. with RTM model

Wentz Correction (Implemented in V5 TMI data product)

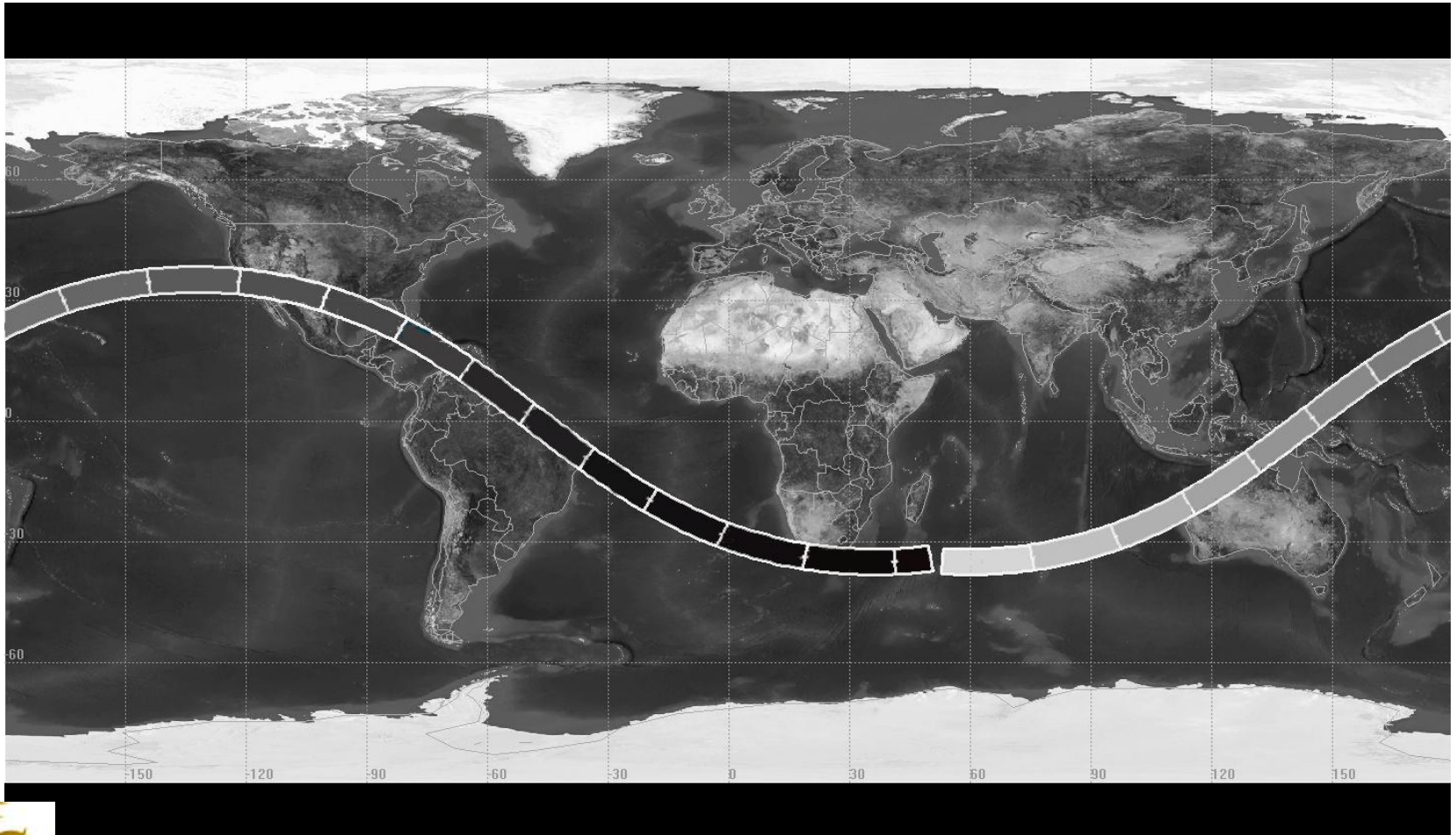
$$T_{ant} = \Gamma T_{sc} + (1 - \Gamma) T_{phy}$$

Constant

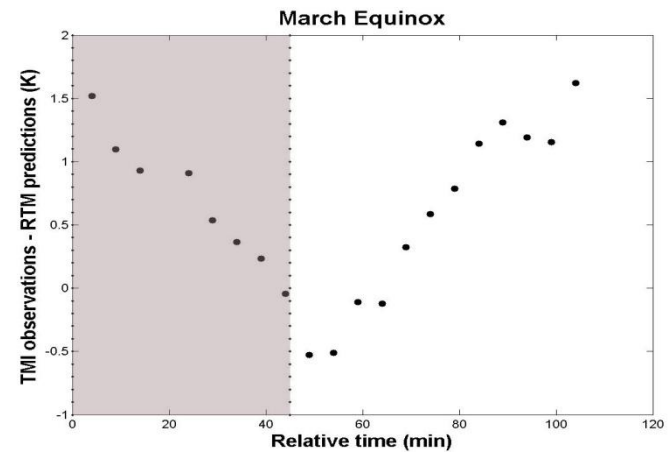
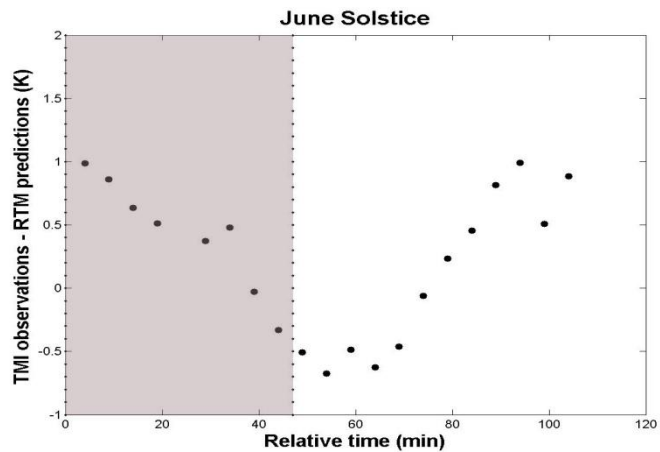
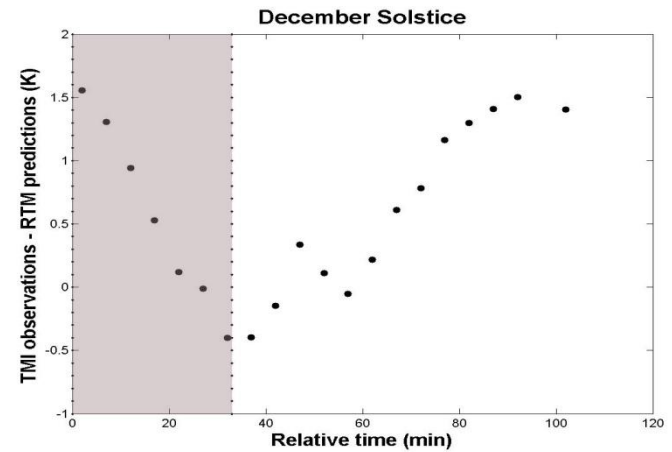
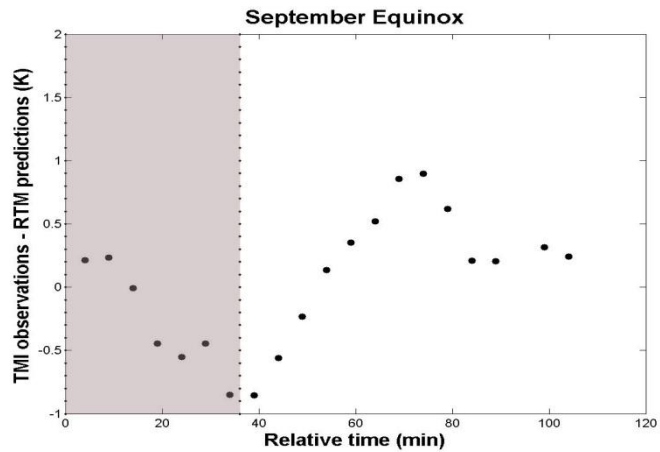
PARAMETERS RELATING TO THE MODEL FOR THE TMI WARM BIAS. THE EMISSIVITY ϵ AND TEMPERATURE T_0 OF THE UNKNOWN EMITTER ARE GIVEN ALONG WITH THE PREDICTED AND OBSERVED WARM BIAS AT $T_A = 2.7$ K

Channel	ϵ	T_0 (K)	$\Delta T_{A, \text{pred}}$ (K)	$\Delta T_{A, \text{obs}}$ (K)
19V	0.0370	302.3	11.1	12.4
19H	0.0284	290.4	8.2	12.3
21V	0.0377	294.6	11.0	13.5
37V	0.0375	296.1	11.0	13.2
37H	0.0274	294.7	8.0	12.2
85V	0.0396	279.6	11.0	13.7
85H	0.0277	239.6	6.6	13.0

5 minute averaging over TMI orbit

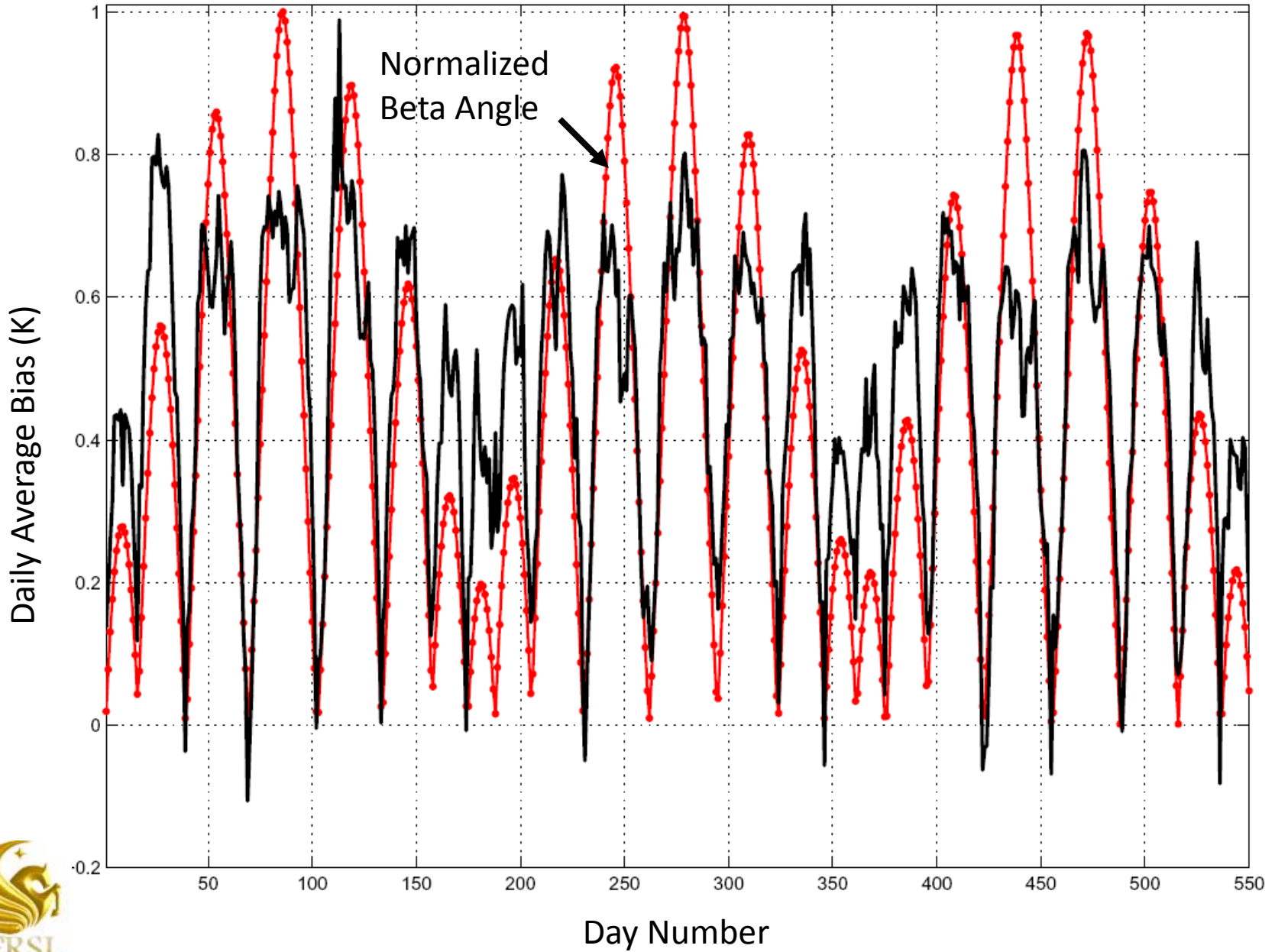


TMI bias variation with orbit time

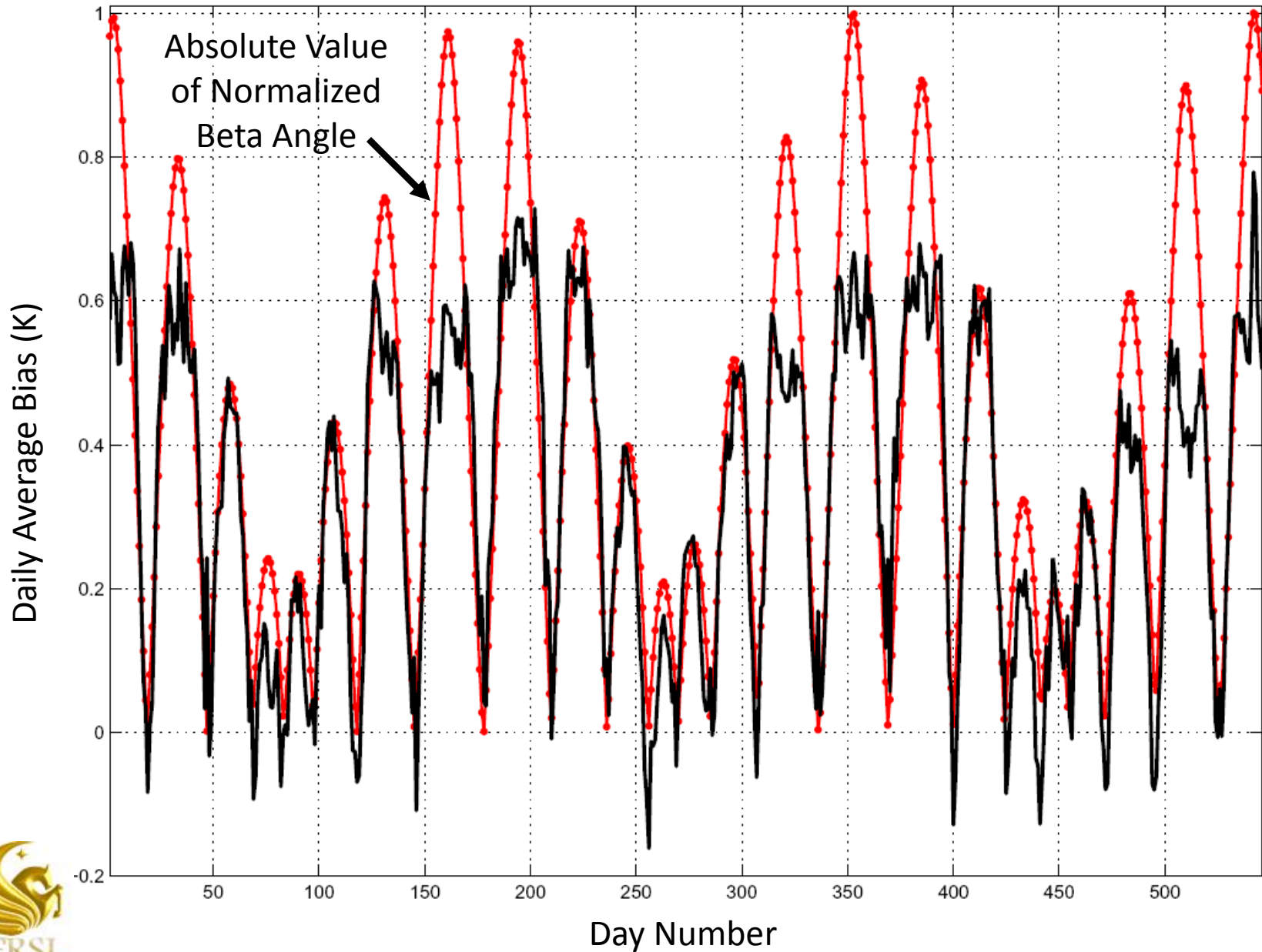


Bias variation with Solar Beta angle

TMI Daily Average Bias Plot For Positive Beta Angles [2004-2006]

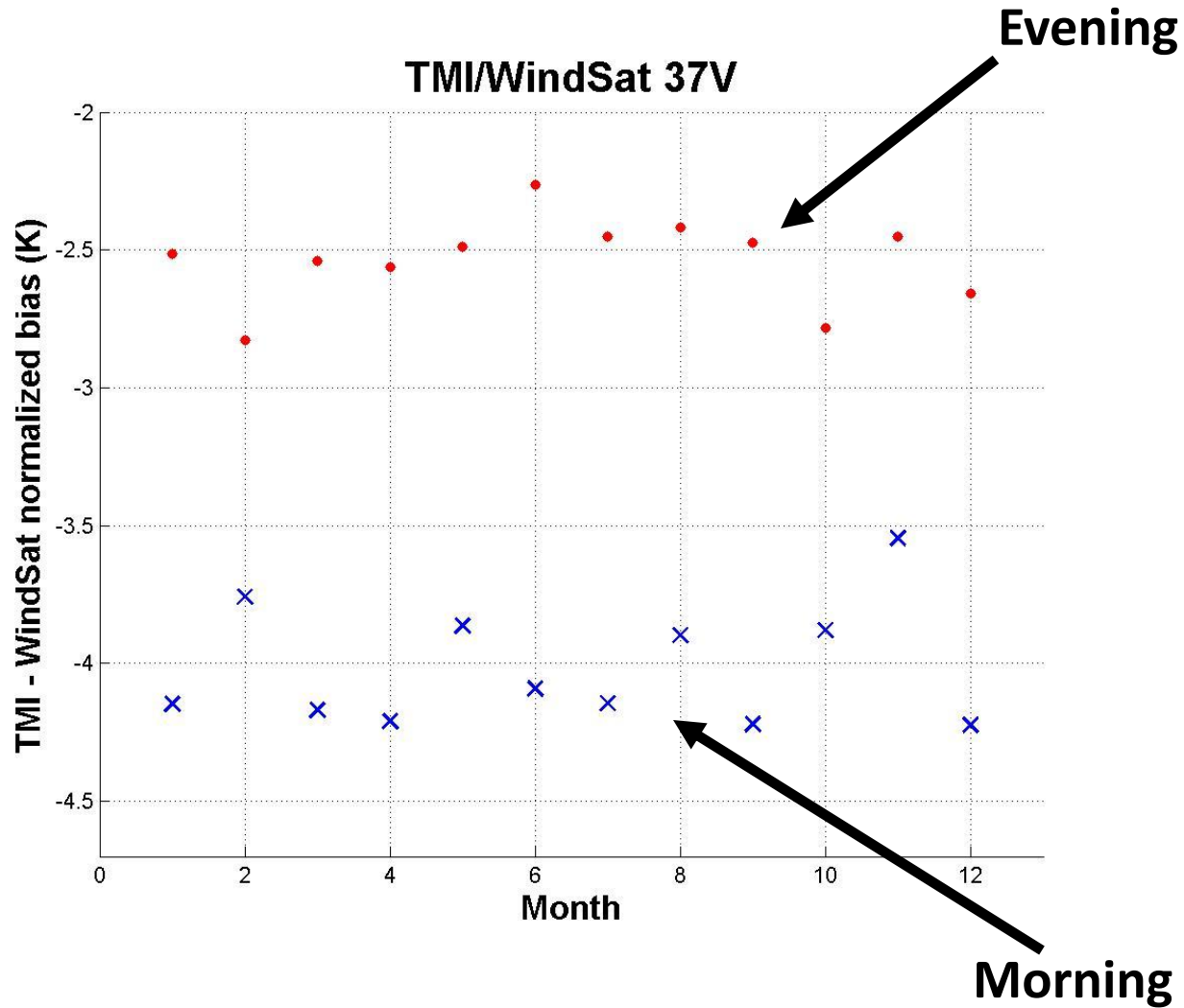


TMI Daily Average Bias Plot For Negative Beta Angles [2004-2006]

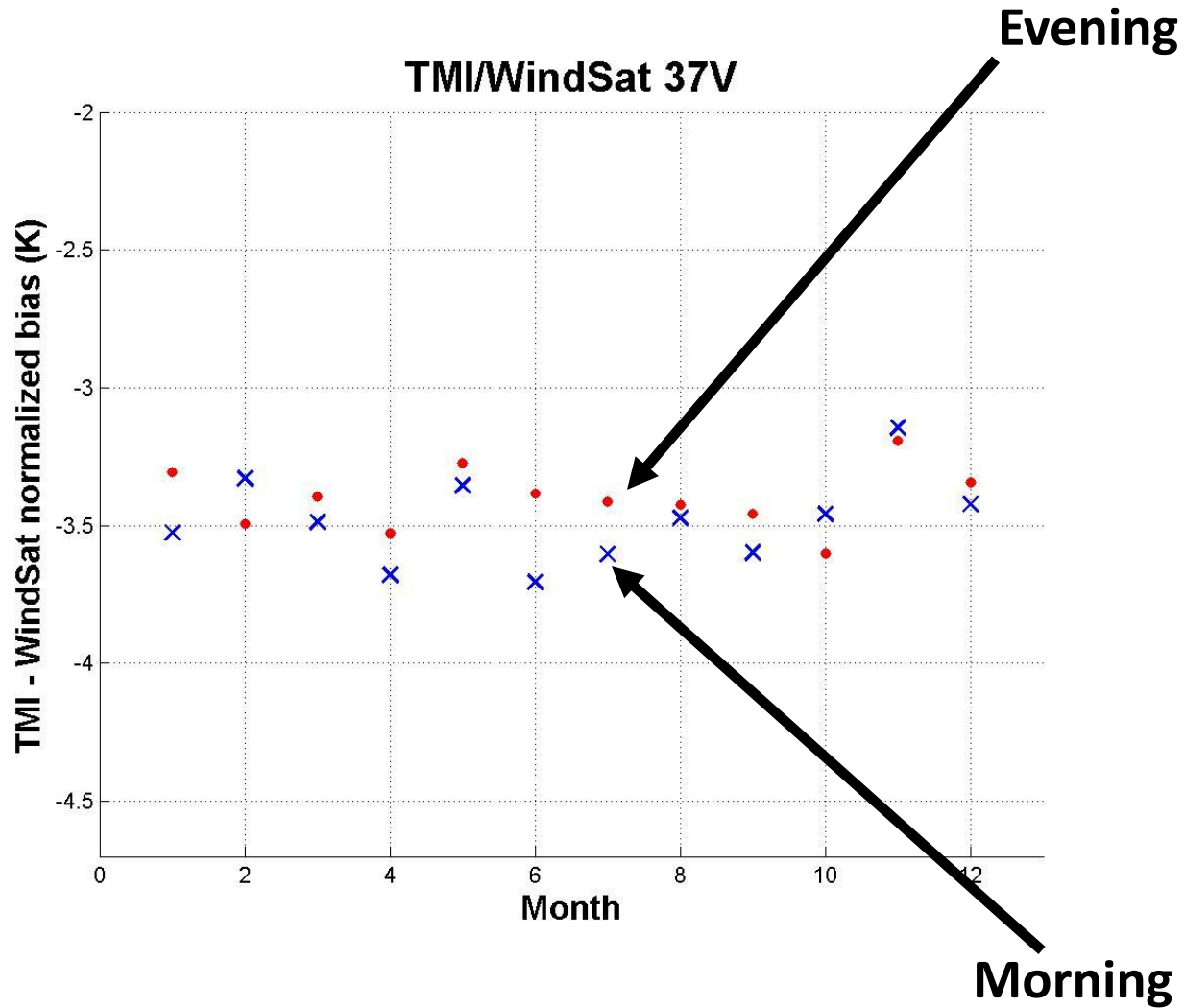


Effect of TMI time-varying bias correction on Inter-Cal results

WindSat Uncorrected



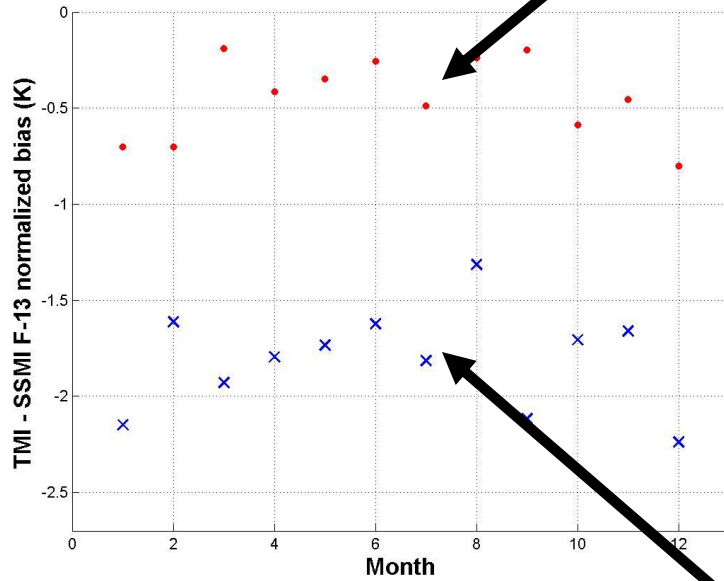
WindSat Corrected



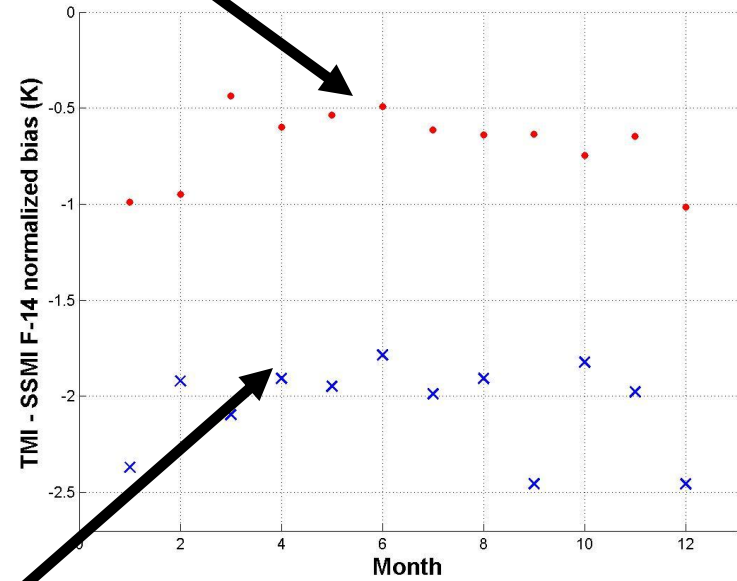
SSMI Uncorrected

Evening

TMI/SSMI F-13 37V



TMI/SSMI F-14 37V

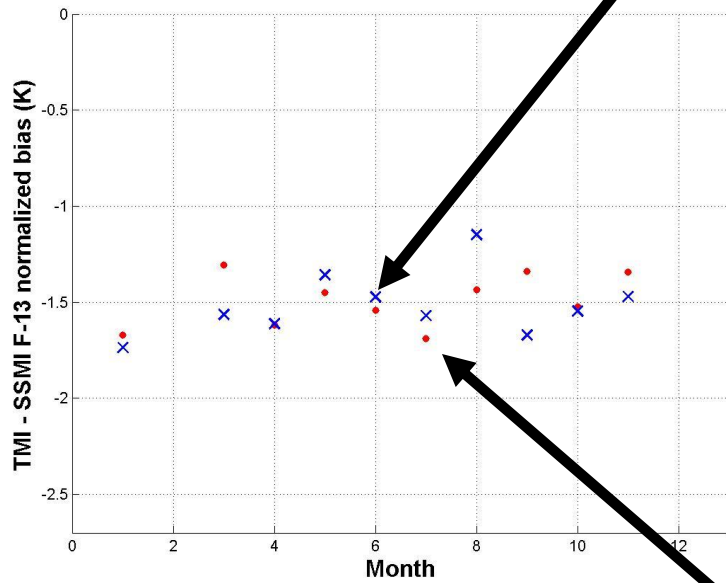


Morning

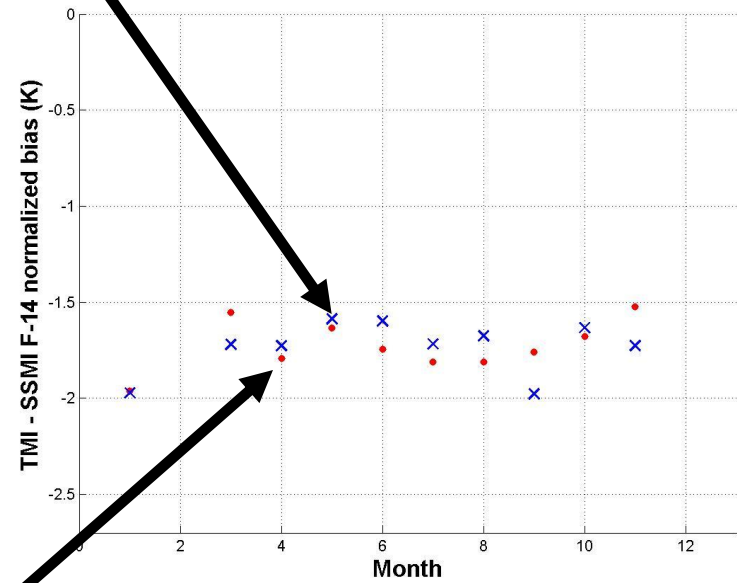
SSMI Corrected

Evening

TMI/SSMI F-13 37V



TMI/SSMI F-14 37V



Morning



SSMI V-pol biases (Uncorrected)

Tb bias	F-13		F-14	
	Evening	Morning	Evening	Morning
19V	-0.44	-1.86	-0.07	-1.43
21V	-1.61	-2.99	-1.23	-2.62
37V	-0.44	-1.82	-0.68	-2.06

SSMI V-pol biases (Corrected)

Tb bias	F-13		F-14	
	Evening	Morning	Evening	Morning
19V	-1.50	-1.58	-1.12	-1.13
21V	-2.67	-2.71	-2.28	-2.32
37V	-1.50	-1.54	-1.73	-1.76

TMI-WindSat V-pol biases

Tb bias	Uncorrected		Corrected	
	Evening	Morning	Evening	Morning
10V	0.87	-0.54	0.00	-0.01
19V	0.16	-1.38	-0.71	-0.85
21V	-0.95	-2.58	-1.82	-2.05
37V	-2.53	-4.01	-3.40	-3.49

V7 Radiometric Bias

- Estimated bias applied in the form of look-up table based on orbit time after eclipse and beta angle
- Beta angle : 1 degree (or finer) quantization
- Orbit time : 1 minute quantization
- Separate look-up tables for pre-boost and post-boost periods

V7 Radiometric Bias

- Post-boost bias estimate table to be created by averaging estimated bias for all 1 degree pixels from 2004-2007 for every beta/orbit time bin
- Pre-boost bias estimate table to be created by averaging estimated bias for all available 1 degree pixels for every beta/orbit time bin
- Smoothing applied to table for both beta and time

Summary

- Radiometric Inter-calibration performed for ICWG dataset
- Time-varying TMI bias discovered, and corrected using an empirical method based on comparing TMI observations to RTM
- TMI error estimated as function of orbit time after eclipse and sun angle
- Correction will be applied to TRMM 1B11 v7, scheduled for release in 2010



Publications

- K. Gopalan, L. Jones, T. Wilheit, and T. Kasparis, “Inter-Satellite Radiometer Calibration of Windsat, TMI AND SSMI”, *IGARSS 2008*.
- K. Gopalan, L. Jones, S. Biswas, S. Bilanow T. Wilheit, and T. Kasparis, “A Time-Varying Radiometric Bias Correction for the TRMM Microwave Imager”, submitted to TGARSS (accepted by GRSL)

Future work

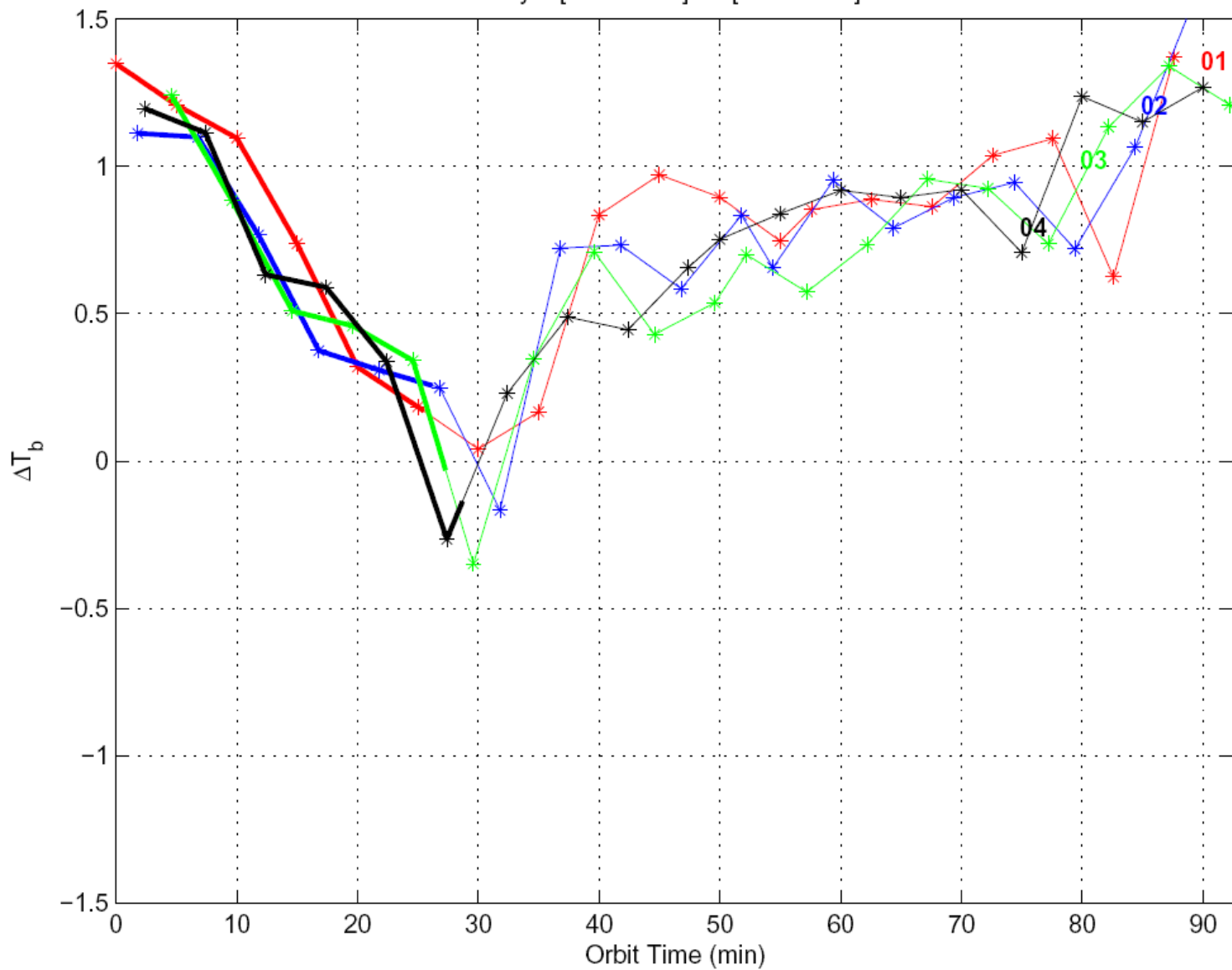
- Analysis of TMI bias for pre-boost period
- Direct comparisons between WindSat and SSMI in polar regions
- Comparison of WindSat and RTM
- Detection of possible TMI pointing errors through comparison with RTM

Back-up slides

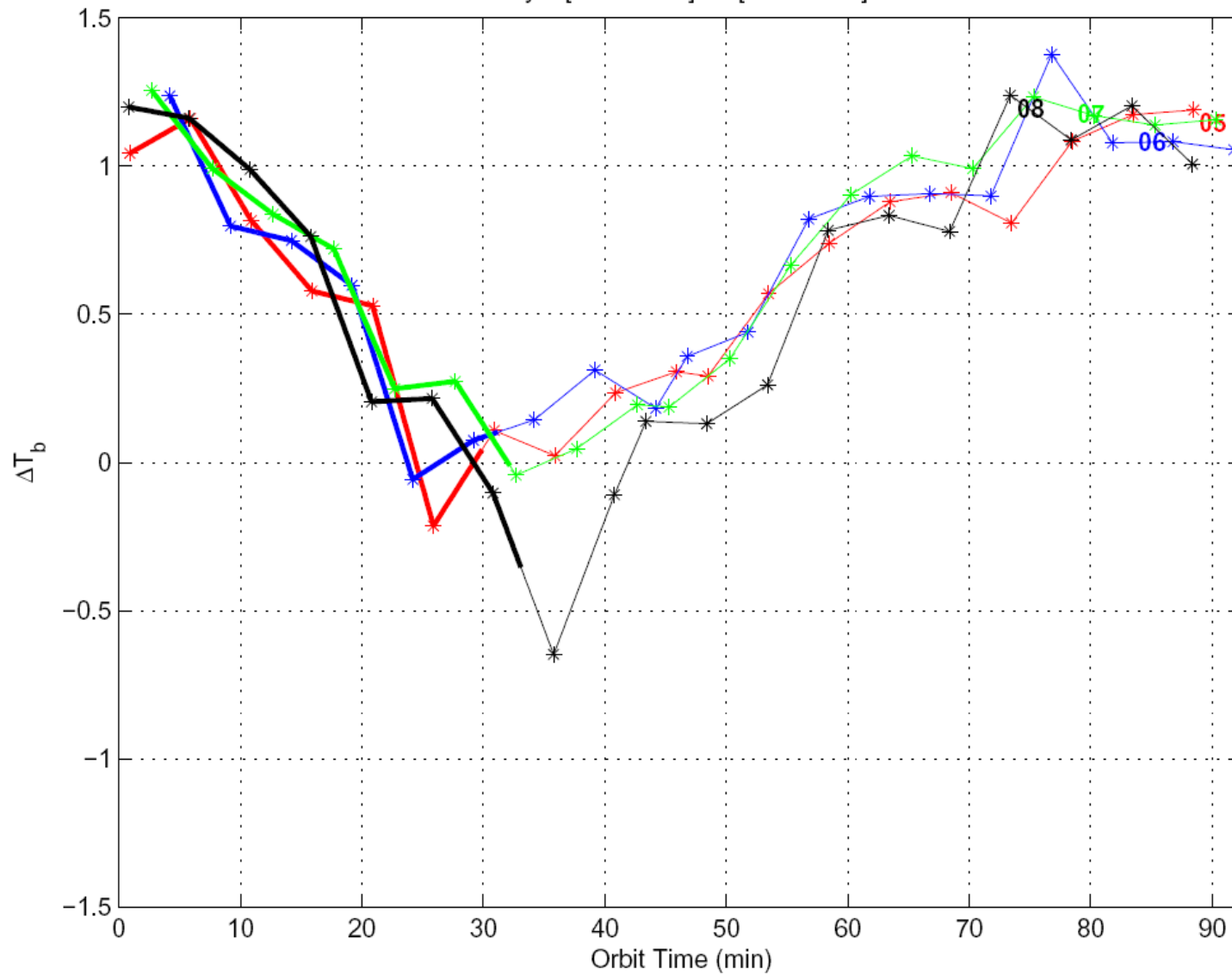
TMI bias dependence on sun angles

- Analysis was primarily performed by Sayak Biswas
- Direct contributions from Linwood Jones, Kaushik Gopalan and Steve Bilanow

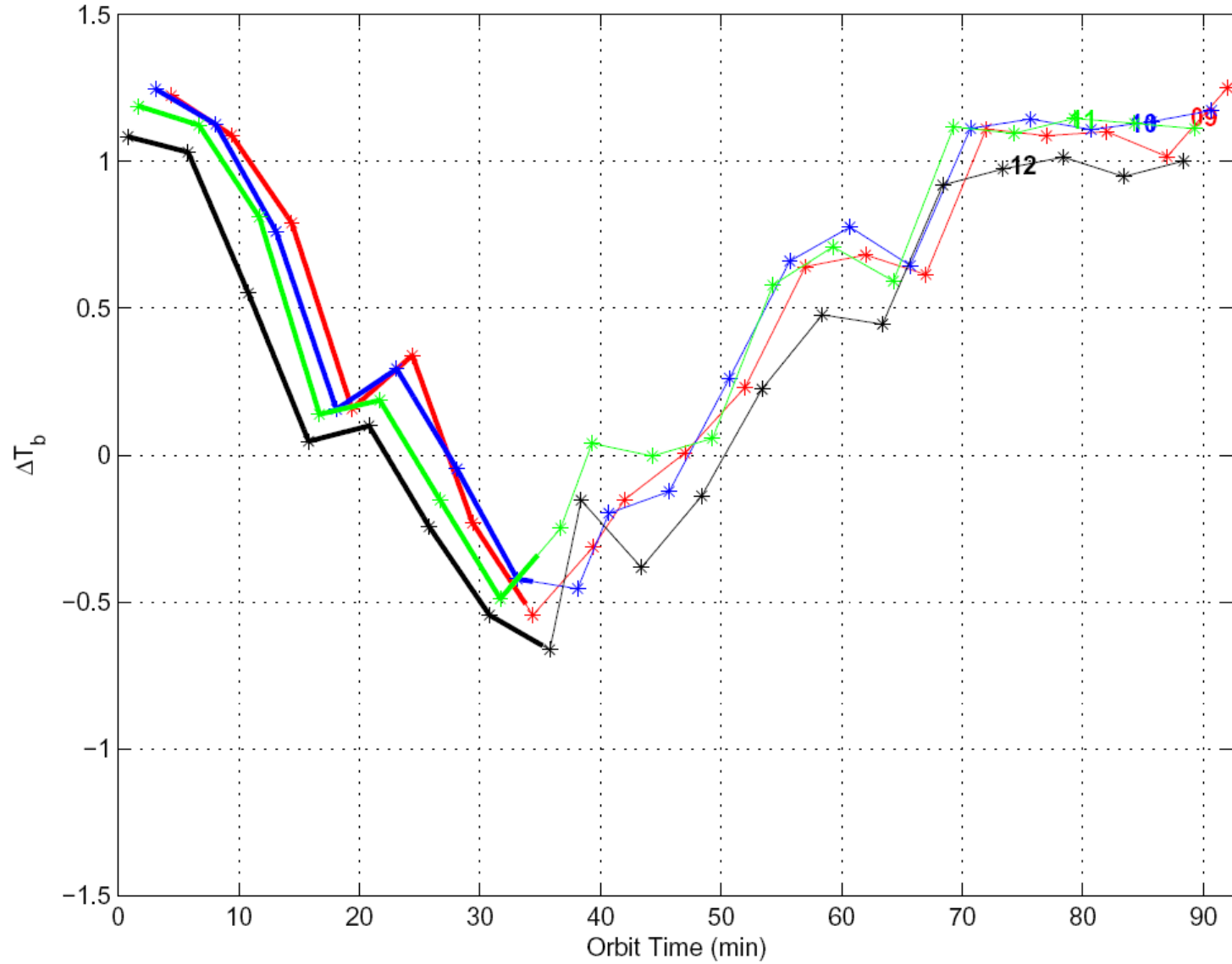
day01[1 Jul 2005]-04[4 Jul 2005]



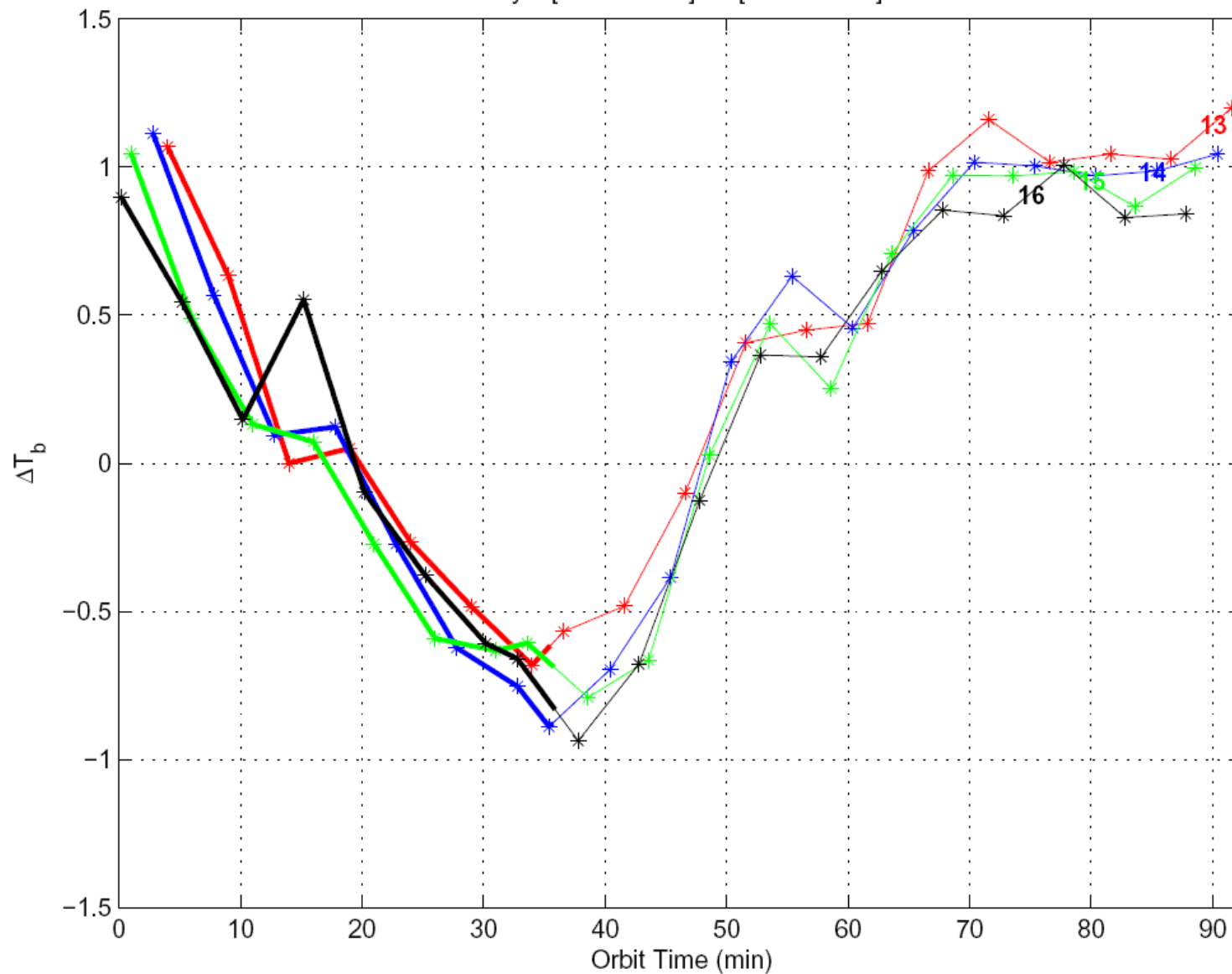
day05[5 Jul 2005]-08[8 Jul 2005]



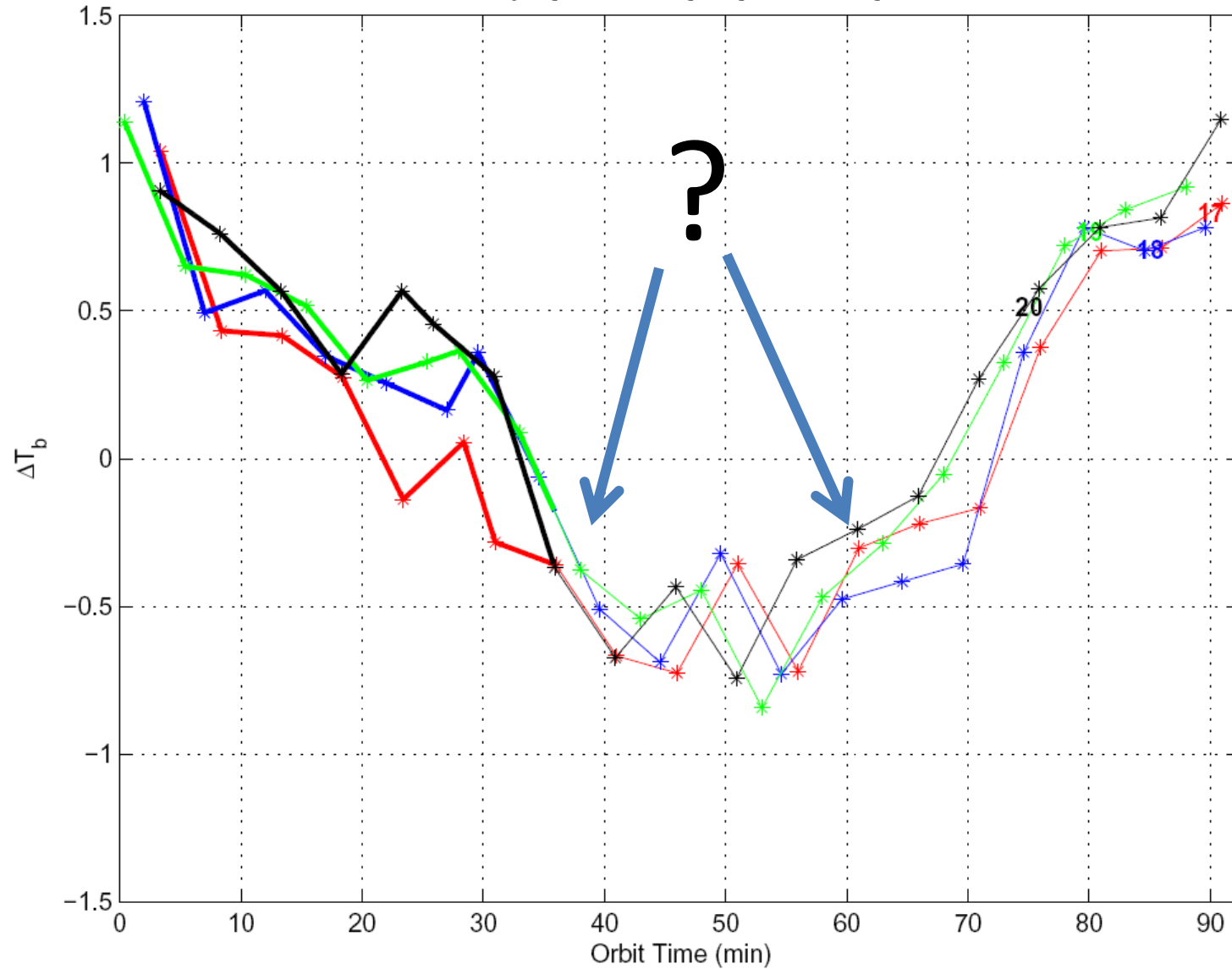
day09[9 Jul 2005]-12[12 Jul 2005]



day13[13 Jul 2005]-16[16 Jul 2005]

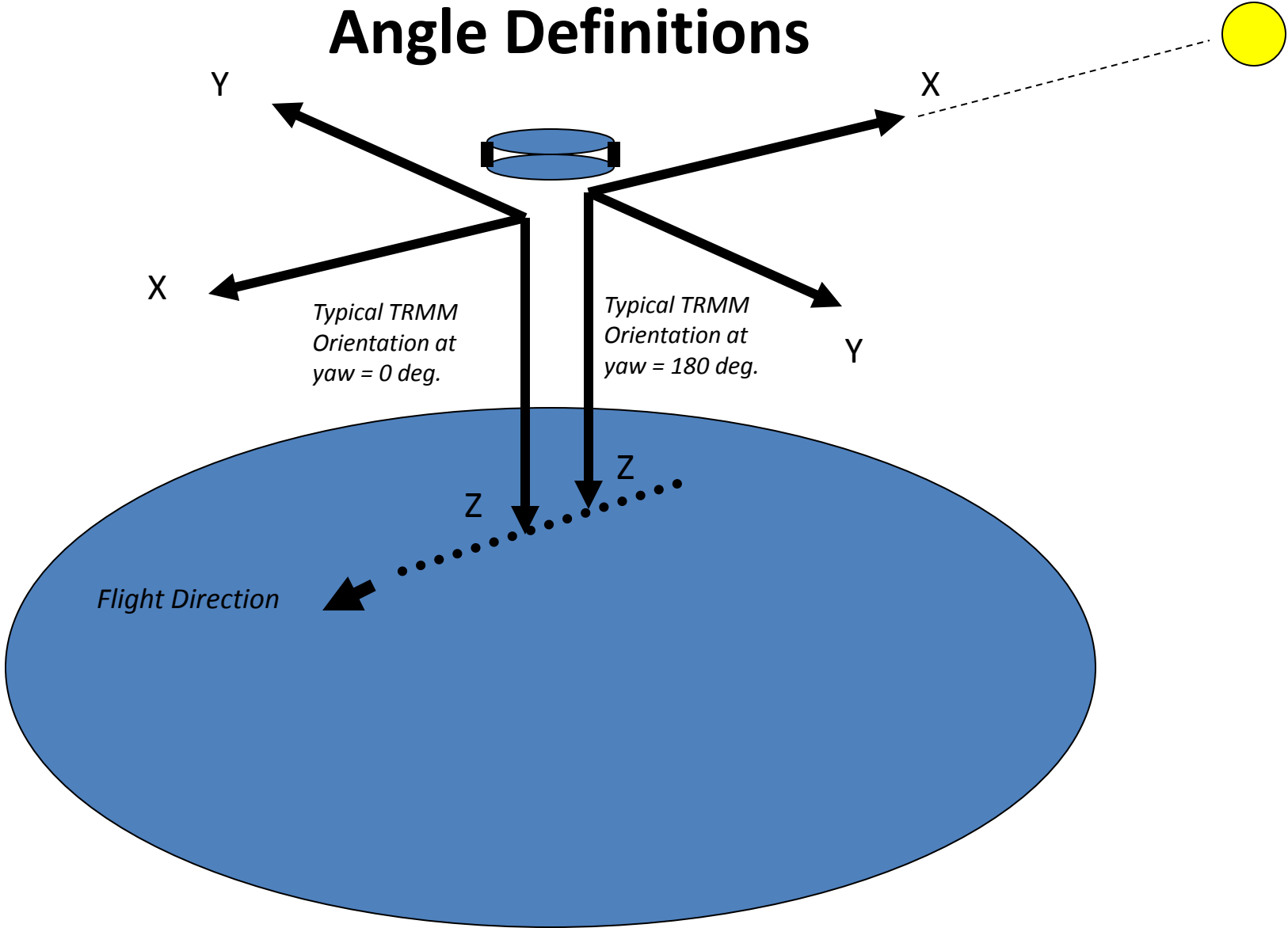


day17[17 Jul 2005]-20[20 Jul 2005]



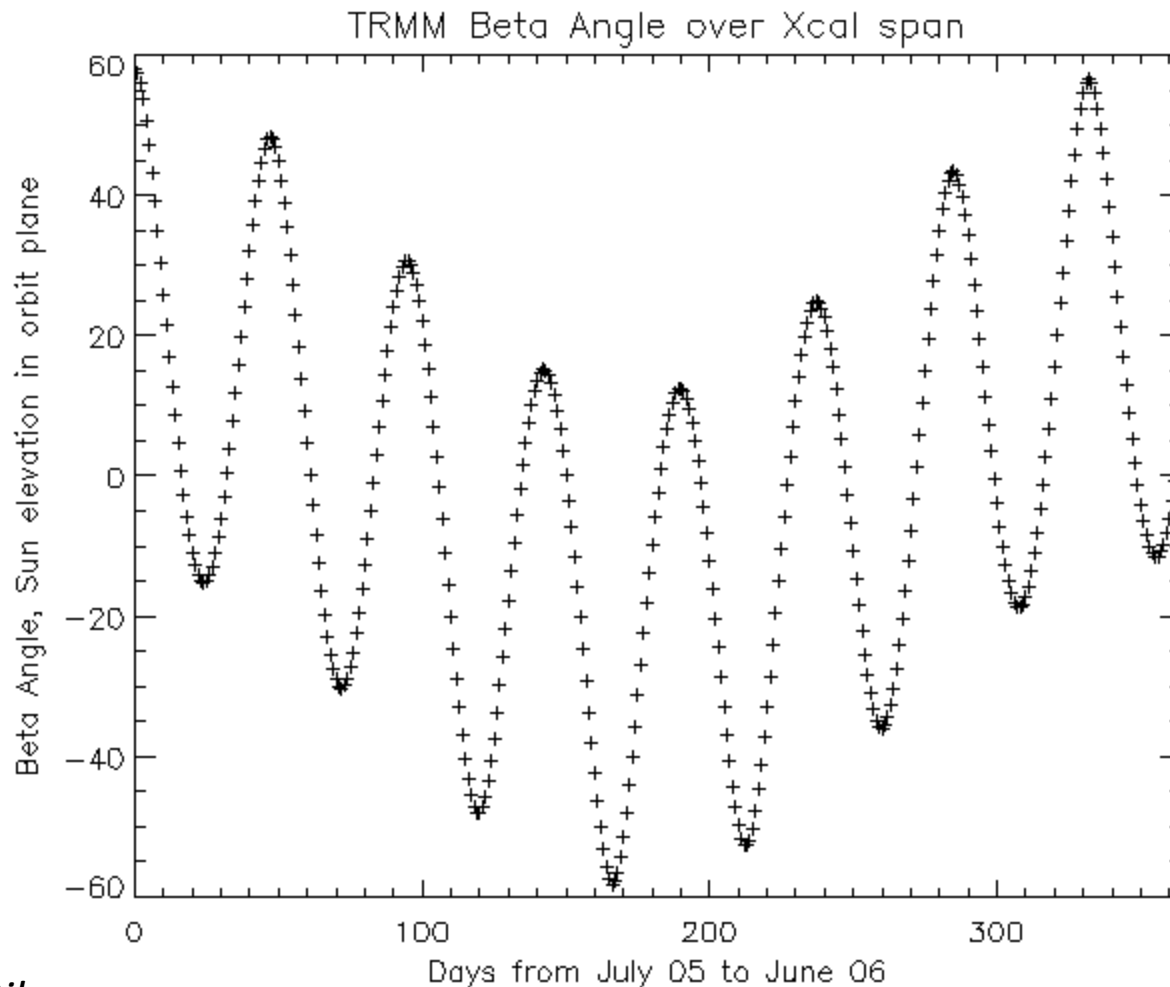
Courtesy: Steve Bilanow

Body or Instrument Coordinate Angle Definitions



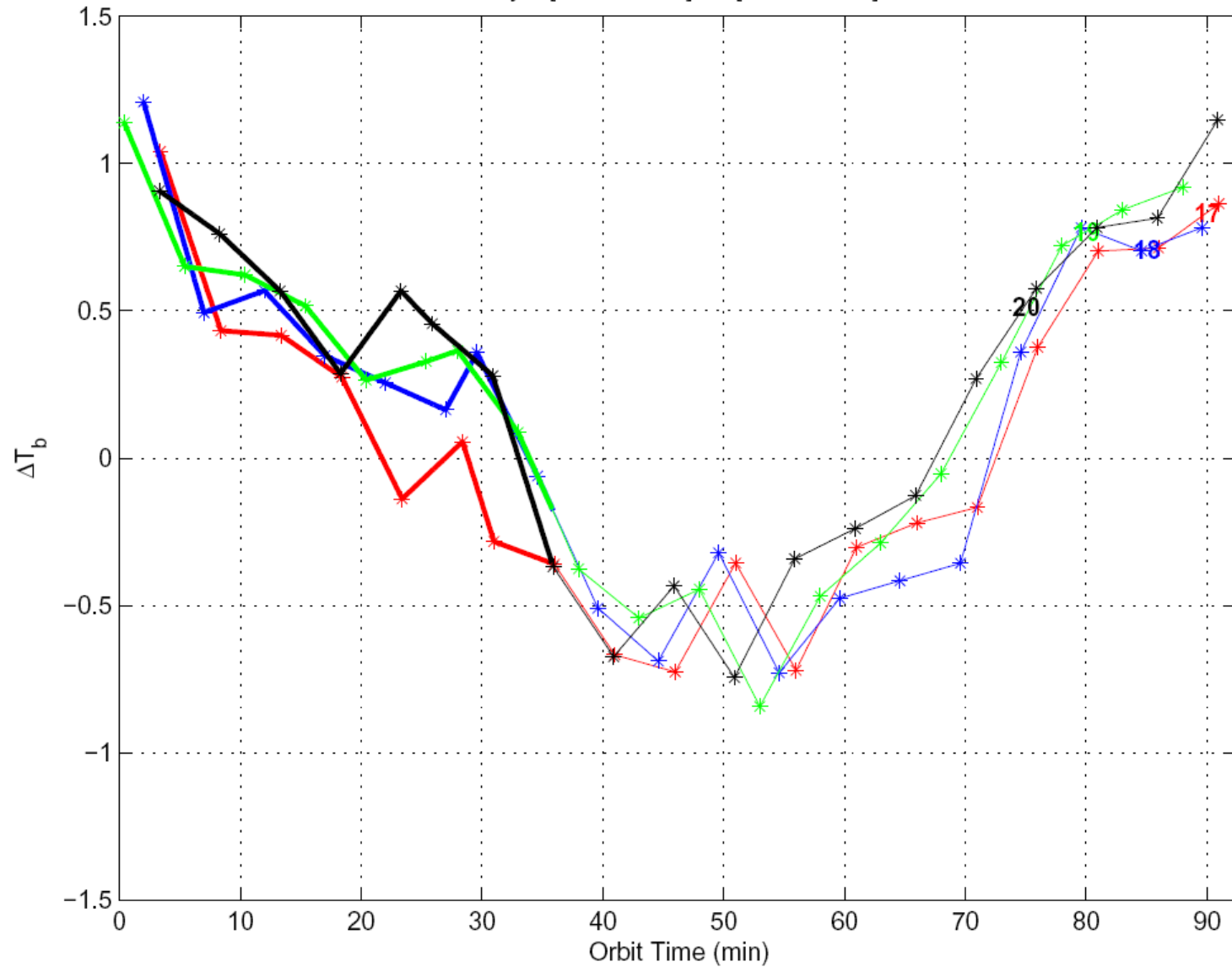
Solar beta angle : Sun elevation above X-Y plane positive toward $-Z$

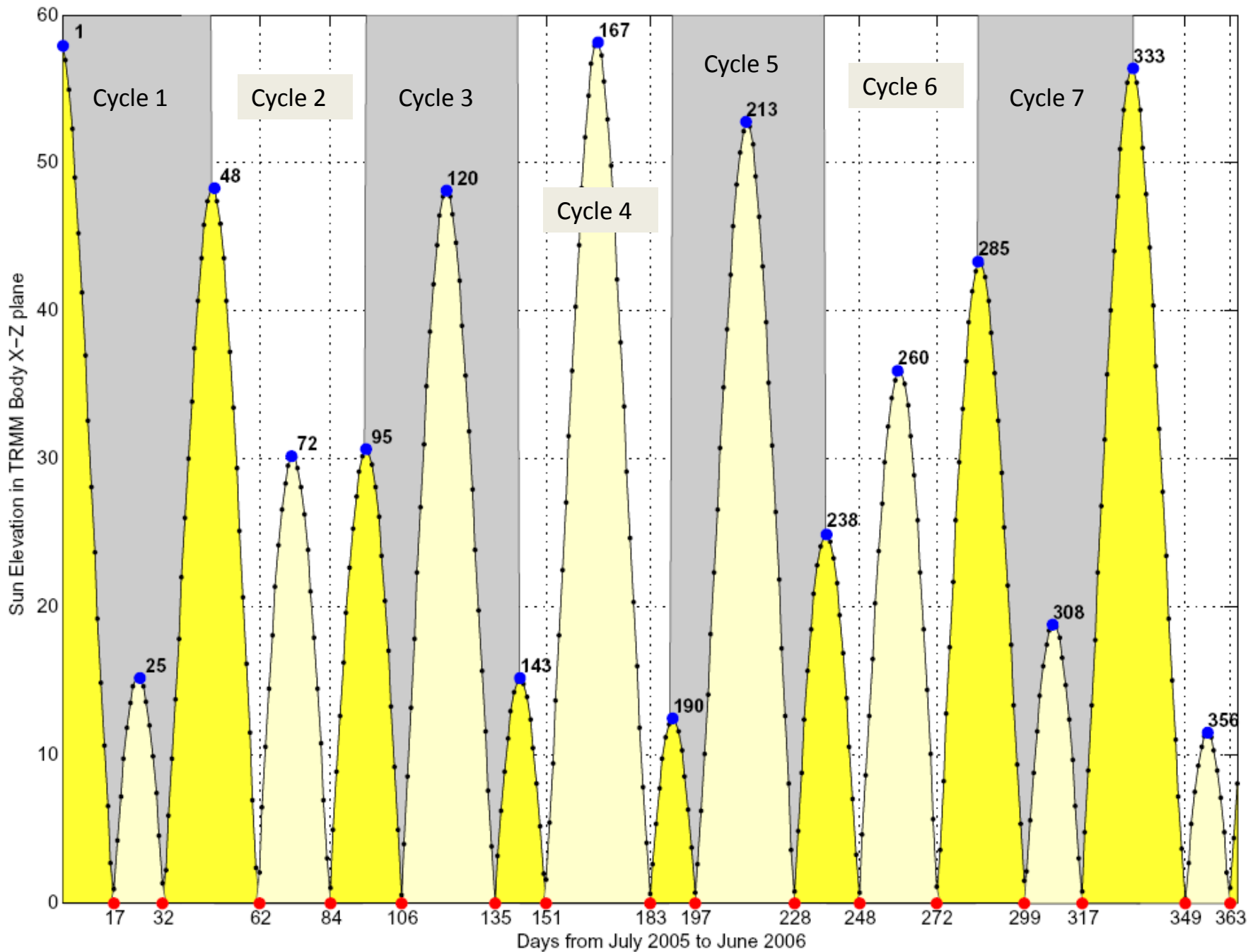
TRMM Solar Beta Angle Variation over 1 year is a combination of +/- 35 degrees due to the orbit inclination on a 46 or 47 day cycle, and the annual variation of +/- 23.5 degrees due to the Earth's axis tilt.



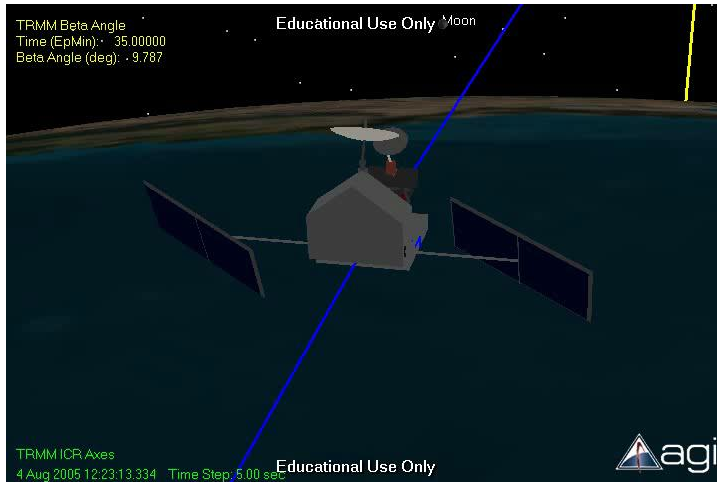
Courtesy: Steve Bilanow

day17[17 Jul 2005]-20[20 Jul 2005]

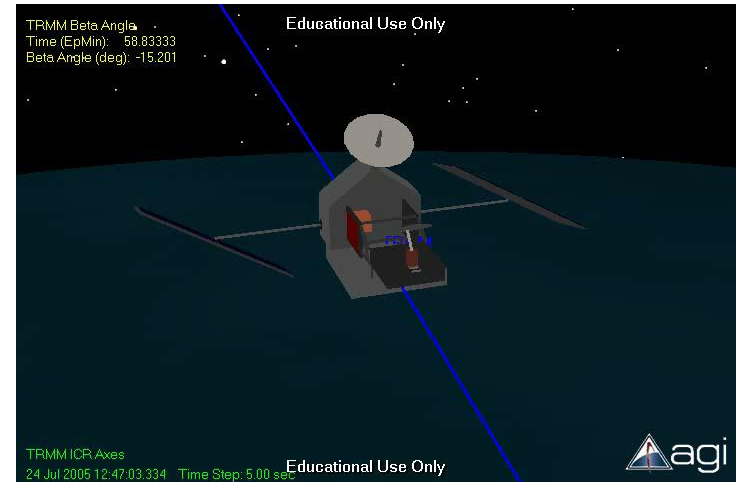




Yaw Flip



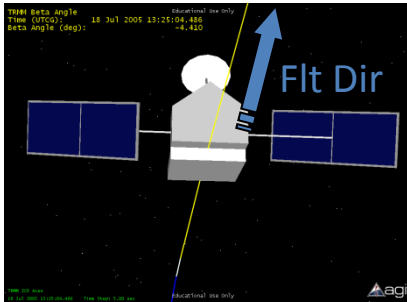
$\text{Beta} > 0^{\circ}$
+X body axis forward
yaw angle = 0°



$\text{Beta} < 0^{\circ}$
-X body axis forward
yaw angle = 180°

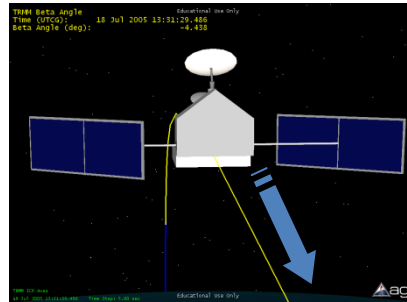
View From The Sun (negative beta angle)

Spacecraft Aft Structure Shadowing of TMI - Jul 18, 2005 [$\beta = -4.4^\circ$]



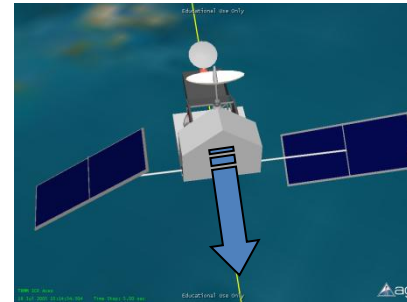
TRMM exits eclipse
(TMI in shadow)

A



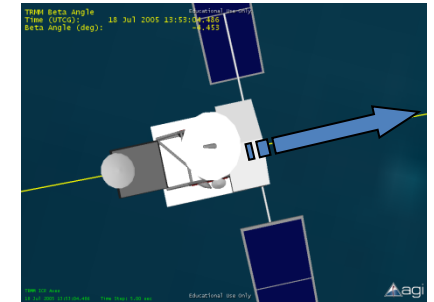
TMI exits S/C shadow

B



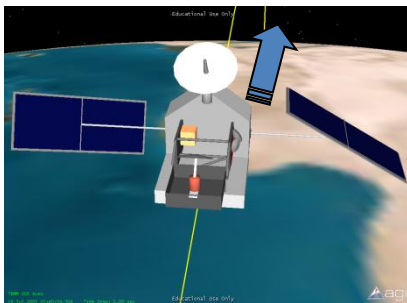
TMI in sunlight

C



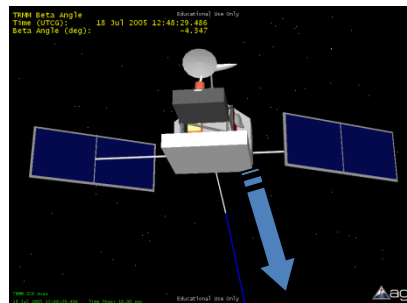
TMI in sunlight

D



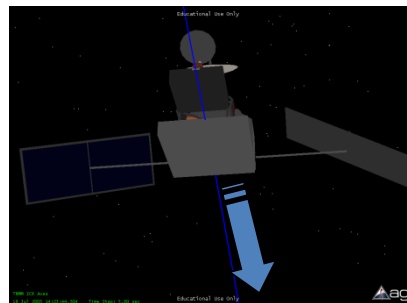
TMI in sunlight

E



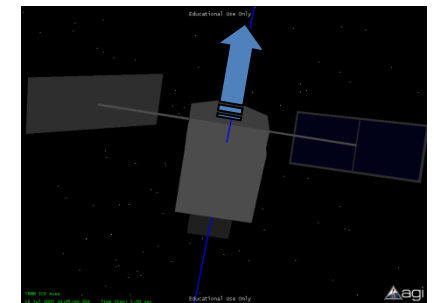
TMI reflector face
illuminated before
eclipse

F



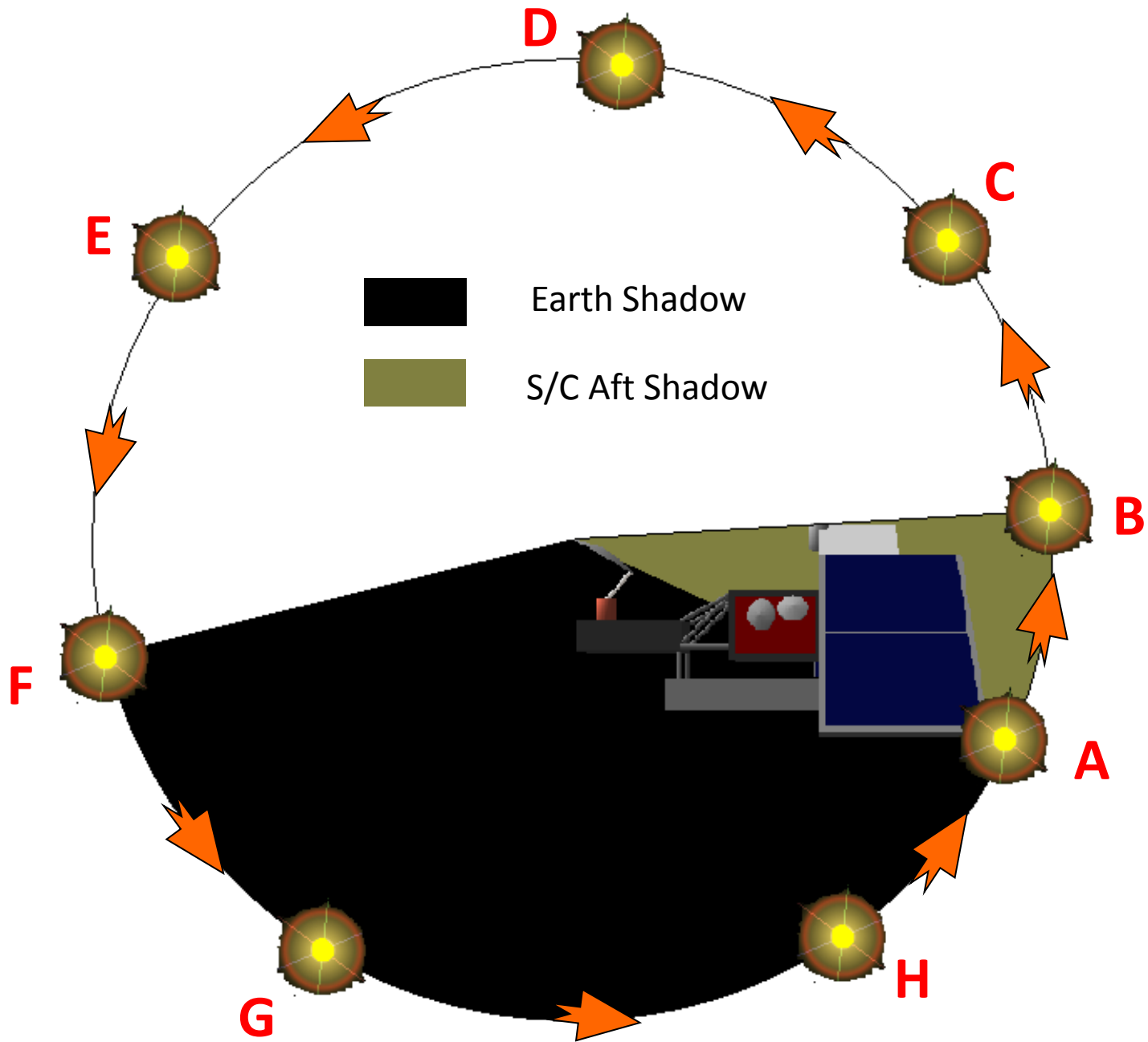
TMI in eclipse

G

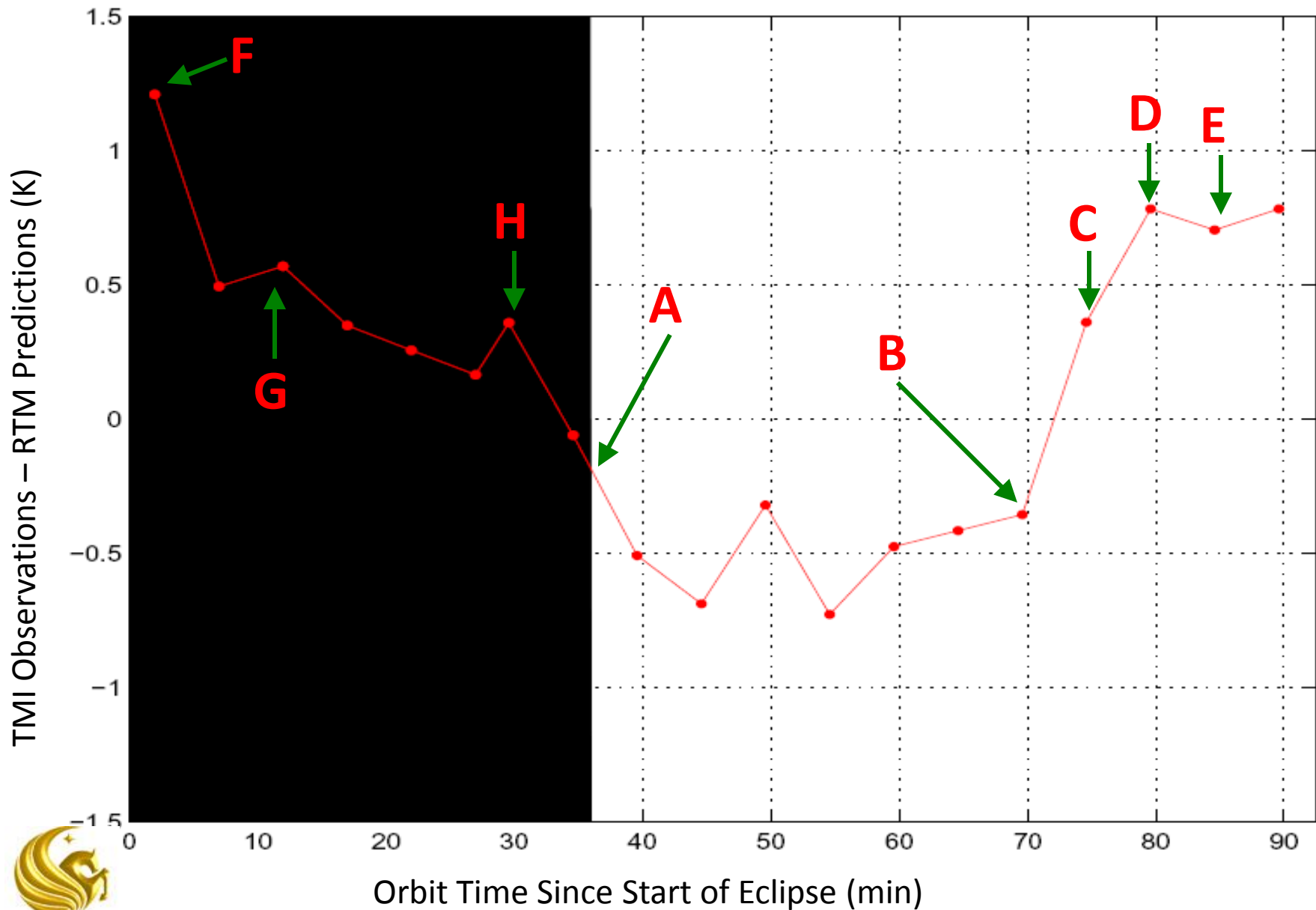


TMI in eclipse

H

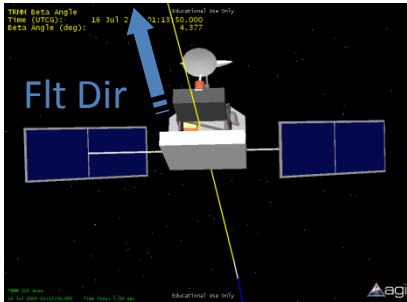


Spacecraft Aft Structure Shadowing of TMI - Jul 18, 2005 [$\beta = -4.4^\circ$]



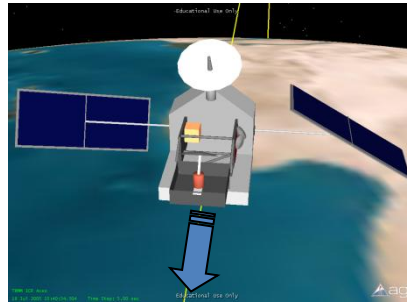
View From The Sun (positive beta angle)

Spacecraft Aft Structure Shadowing of TMI – Aug 02, 2005 [$\beta = +4.3^\circ$]



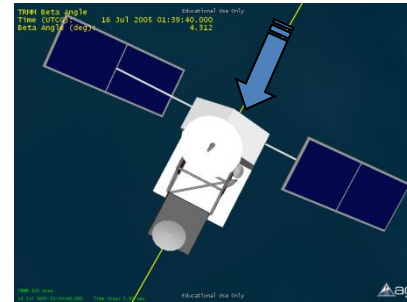
TRMM exits eclipse
(TMI in sunlight)

A



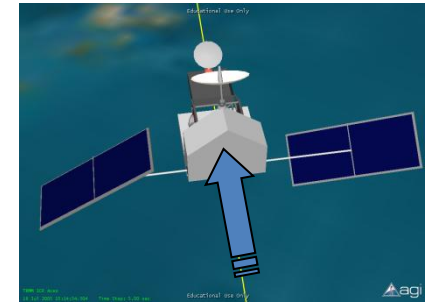
TMI in sunlight

B



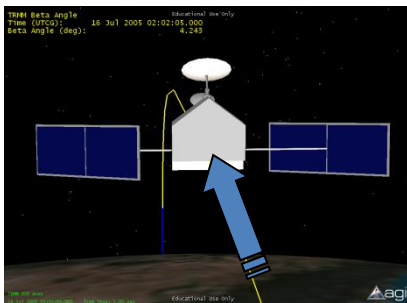
TMI in sunlight

C



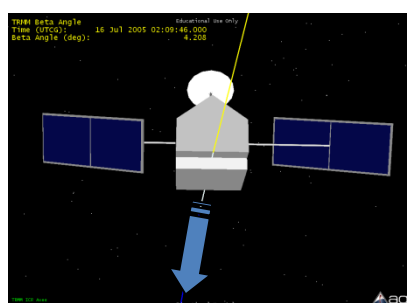
TMI in sunlight

D



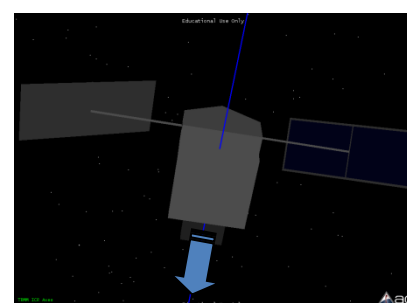
TMI enters S/C shadow

E



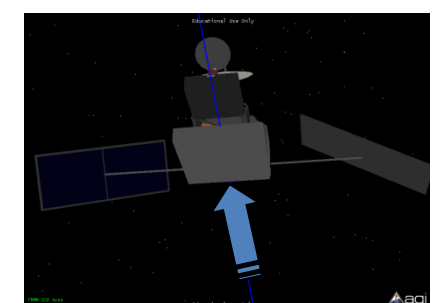
TMI in S/C shadow before
eclipse

F



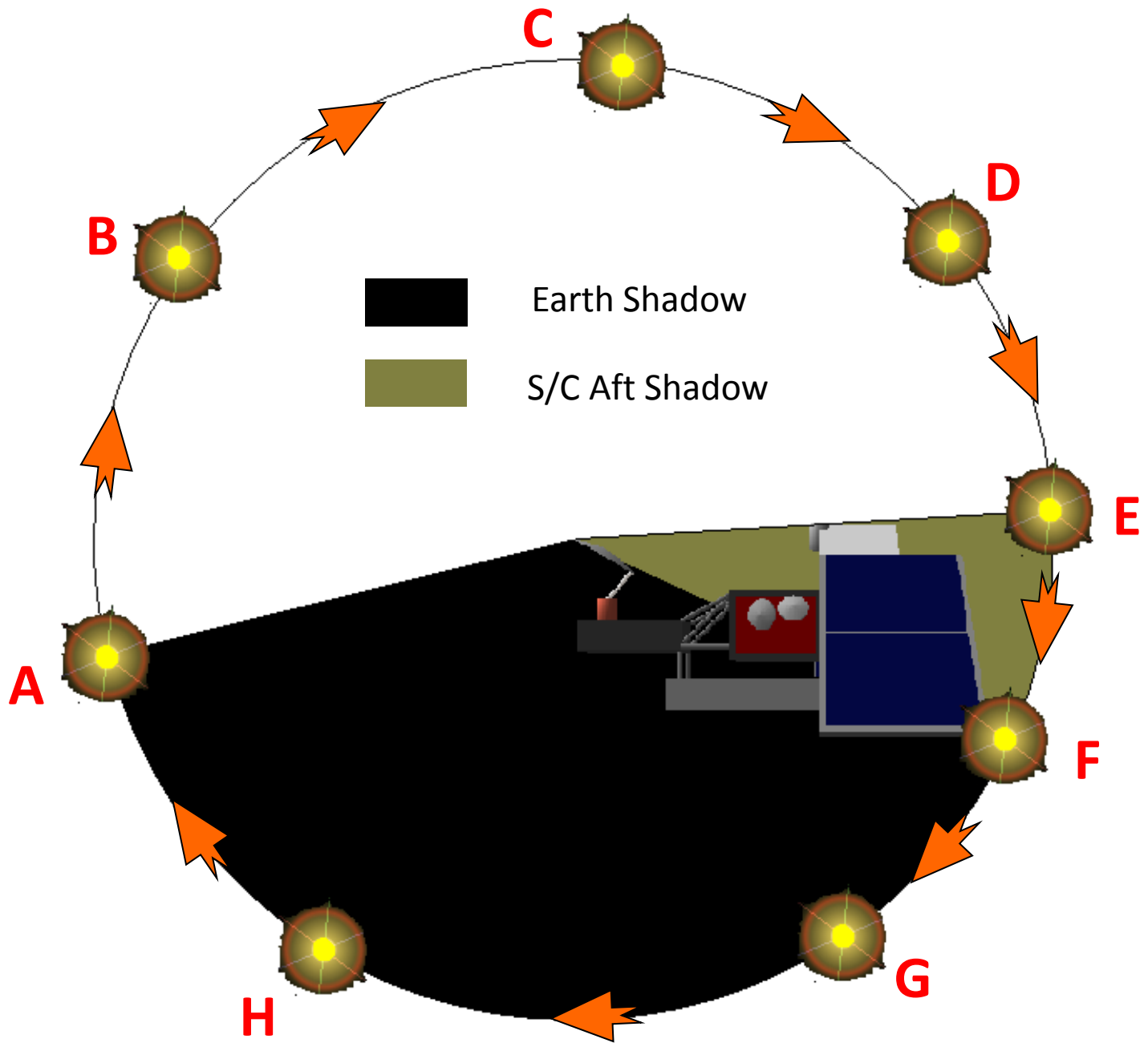
TMI in eclipse

G



TMI in eclipse

H



Spacecraft Aft Structure Shadowing of TMI – Aug 02, 2005 [$\beta = +4.3^\circ$]

