



Simulation and Study of the Stokes Vector in a Precipitating Atmosphere

Ian Stuart Adams Doctoral Dissertation Final Examination 29 March 2007







- Part I: Background
 - Passive Microwave Remote Sensing
 - Wave Polarization
 - Scattering Theory and Radiative Transfer
 - The T-Matrix
- Part II: Radiative Transfer Model
 - Precipitation Profiles
 - Gas Absorption
 - Scattering Calculations
- Part III: Results





Part I





- Passively measure microwave energy
 - Absorption, emission, and reflection
- Application in geophysical retrievals
 - e.g., Precipitation, wind
- Use range of frequencies
 - Simultaneously solve interdependent variables
- Strong heritage
 - SMMR, SSM/I
- WindSat
 - 6.8, 23.8 GHz v & h
 - 10.7, 18.7, 37.0 GHz fully polarimetric





- Gas
 - Unpolarized absorption/emission in microwave region
 - Primarily water vapor
- Hydrometeors
 - Non-spherical particles
 - Absorption/emission and scattering frequency dependent
 - Polarizing
- Sea surface temperature
 - Low frequency effects
- Ocean surface wind vector
 - Speed and direction dependence
 - Small directional signal
 - Polarized emission and reflection



Precipitation Interference



- Emission
 - Large compared other ocean parameters
 - Polarized
- Scattering
 - Reduces measured surface emission signal
 - Alters incident polarization
- Absorption
 - Related to emission
 - Reduces measured surface emission signal
- Surface perturbations
 - Changes small scale surface structure
 - Little information available



Geophysical Models



Empirical models

- Requires large sets of training data
- Only conditioned for cases included in training data
 - Can lack physical insight
- Relatively simple to implement
- Physics-based models
 - Broad application
 - More difficult to implement





- Orientation of electric (and magnetic) fields
- Effected by particle shape or surface structure
 - Drop shape
 - Wind direction
- Defined by Stokes vector
- Polarization
- Orientation
- Ellipticity





$$\mathbf{I} = \begin{bmatrix} I \\ Q \\ U \\ V \end{bmatrix} = \frac{1}{2}\sqrt{\frac{\epsilon}{\mu}} \begin{bmatrix} E_{0v}E_{0v}^* + E_{0h}E_{0h}^* \\ E_{0v}E_{0v}^* - E_{0h}E_{0h}^* \\ -E_{0v}E_{0h}^* - E_{0h}E_{0v}^* \\ j(E_{0h}E_{0v}^* - E_{0v}E_{0h}^*) \end{bmatrix} = \frac{1}{2}\sqrt{\frac{\epsilon}{\mu}} \begin{bmatrix} E_{0v}E_{0v}^* + E_{0h}E_{0h}^* \\ E_{0v}E_{0v}^* - E_{0h}E_{0h}^* \\ -2\operatorname{Re}(E_{0v}E_{0h}^*) \\ 2\operatorname{Im}(E_{0v}E_{0h}^*) \end{bmatrix}$$



Coordinate System







Polarization Ellipse





Orientation

$$\tan 2\gamma = -\frac{U}{Q}$$

Ellipticity

$$\tan 2\psi = -\frac{V}{\sqrt{Q^2+U^2}}.$$





Amplitude scattering matrix

$$\begin{bmatrix} E_v^s(r\hat{\mathbf{n}}) \\ E_h^s(r\hat{\mathbf{n}}) \end{bmatrix} = \frac{\mathrm{e}^{\mathrm{j}k_1r}}{r} \begin{bmatrix} S_{11} & S_{12} \\ S_{21} & S_{22} \end{bmatrix} \begin{bmatrix} E_{0v}^i \\ E_{0h}^i \end{bmatrix}$$

- 4 x 4 scattering matrix
 - Intensity and polarization of scattered wave
- 4 x 4 extinction matrix
 - Intensity and polarization loss of wave
- Absorption vector
 - Difference of extinction and scattering





- Emission due to heat
 - All real bodies absorb and emit
- Emission of body directly related to absorption
 - Body temperature times emissivity
 - Interface between two media







- Describes flow of energy through media
- Collection of particles considered homogeneous medium
- Includes emission, extinction and scattering

$$\begin{split} \frac{\mathrm{d}}{\mathrm{d}s} \mathbf{I}(\mathbf{r}, \hat{\mathbf{n}}, \omega) &= -N_0(\mathbf{r}) \left\langle \mathbf{K}(\mathbf{r}, \hat{\mathbf{n}}, \omega) \right\rangle \mathbf{I}(\mathbf{r}, \hat{\mathbf{n}}, \omega) \quad \text{extinction} \\ &+ N_0(\mathbf{r}) \left\langle \mathbf{K}_{\mathbf{a}}(\mathbf{r}, \hat{\mathbf{n}}, \omega) \right\rangle I_{T_b}(T, \omega) \quad \text{emission} \\ &+ N_0(\mathbf{r}) \int_{4\pi} \mathrm{d}\hat{\mathbf{n}}' \left\langle \mathbf{Z}(\mathbf{r}, \hat{\mathbf{n}}, \hat{\mathbf{n}}', \omega) \right\rangle \mathbf{I}(\mathbf{r}, \hat{\mathbf{n}}, \omega) \quad \text{scattering} \end{split}$$







- Analytical relationship between incident and scattered energy
- Solved using extended boundary condition method
 - Internal field excited by incident radiation
 - Scattered field resultant of internal field
- Ideally suited for rotationally symmetric particles
- Efficient FORTRAN algorithms developed by Michael Mishchenko





- Isotropic, mirror symmetric medium
- Extinction matrix diagonal
 - No polarization
- Scattering matrix block diagonal
 - I and Q separate from U and V
- Scalar emission
- Spherical particles
 - Special case
 - Diagonal scattering matrix





- Not isotropic
- Some symmetries
 - No azimuthal dependence
- Extinction block diagonal
- I and Q separate from U and V along incidence direction
- Off-angle scattering breaks separation between I/ Q and U/V
- Polarized emission





Part II





- Simulates evolution of convective systems
 - 3 spatial dimensions
 - Time evolution
- Water content profiles
 - Rain
 - Snow
 - Graupel
 - Cloud Liquid
 - Cloud Ice
- Temperature, pressure and humidity information
- Does not represent melting layer
- TOGA-COARE array event
 - 22 February 1993
- Courtesy of Chris Kummerow and Jody Crook



3D Rain Profiles







3D Graupel Profiles







3D Snow Profiles









- Convert water content to particle number density
 - Falling hydrometeors described by Marshall-Palmer type (inverse exponential) distribution

$$N(D) = N_0 \mathrm{e}^{-\lambda D}$$

where

D = Drop diameter

$$N_0 = \text{Intercept}$$

$$\lambda = \text{Slope}$$

N(D) = Number density over D + dD

- Modified gamma used for cloud liquid
- McFarquhar and Heymsfield for cloud ice



Aspect Ratio



- Rain drops flatten due to drag
 - Closely approximated by oblate spheroid
 - Horizontally Aligned
- Snow particle shape related to size and formation
 - Oblate spheroid "1st order" approximation
 - Horizontally aligned in average
 - Intricate structure << incident microwave wavelength

$$\frac{1}{R_A} = c_0 + c_1 D - c_2 D^2 + c_3 D^3 - c_4 D^4$$

D =Equivalent-volume-sphere radius



Complex Permittivity



- Liquid water
 - Double-Debye equation
 - Empirical fit
 - Real and Imaginary frequency and temperature dependent
- Ice
 - Empirical Fit
 - Real constant
 - Imaginary frequency and temperature dependent
 - Maxwell-Garnett mixing for inclusion of air





- Mishchenko T-matrix code
 - Public domain scattering code
- Computationally efficient
- Aspect ratio limited
 - Computational accuracy
 - Only able to process in double precision
 - Particle size
 - Limit for rain beyond largest physical drop size
 - Imaginary component of refractive index
- Separate codes for randomly oriented and aligned particles
 - Spheres considered randomly oriented particles





- Water vapor line at 22.235 GHz
 - Noticeable effect at 18.7 GHz
 - Small at 10.7 and 37 GHz
 - Rosencranz PWR 98 Model
 - Present in trace amounts (< 3%)
- Oxygen continuum
 - Second most abundant constituent of air (~21%)
 - Most noticeable at 37 GHz
 - Small at 18.7 GHz
 - Rosencranz PWR 93 Model
- Nitrogen
 - Negligible at simulated frequencies
 - Most abundant constituent (~78%)





- Ratio of constituent partial pressure to air
- Constant for nitrogen (0.781) and oxygen (0.209)
- Related to water vapor saturation pressure (Flatau) and humidity

$$p_s = a_0 + a_1 T + a_2 T^2 + a_3 T^3 + a_4 T^4 + a_5 T^5 + a_6 T^6$$

where

$$T$$
 = Temperature [Celcius]

$$p_{H_2O} = p_s \times RH$$







- Radiative transfer calculations
 - 3D curved atmosphere
 - Scattering
- Separate gas absorption calculations
 - Generate lookup table
 - Does not account for varying humidity
 - Must generate for each profile
- Monte Carlo scattering
- Customizable surface
- User defined instrument characteristics
- Open source







- Photon tracing from instrument to photon origin
 - Random path lengths and directions
 - Calculate contributions
 - Across path
 - At end of path step
 - Decision at end of path step
 - Scatter (repeat random path length and direction)
 - Absorb (start with new photon)
- Only in specified cloud box
- 3D capable
- Accuracy depends on number of photons
 - High water content requires more photons







• Python interface to ARTS

- General control of ARTS and ARTS scattering
- Template for ARTS handling of scattering parameters
- Contains distributions for clouds
- Includes Mishchenko's T-matrix FORTRAN code
 - Compiled as Python module
- Open source





Part III



Related Work



- Roberti and Kummerow
 - Polarization due to precipitation
 - Primarily V and H (I and Q)
 - Small 3rd Stokes
 - No data
- Kutuza et al.
 - Canted rain drops
 - Simplistic model
 - Large third stokes



Simulation Data



- Simulation Set 1
 - 2D Slice of 3D profile
 - Spatially correlated data set
 - Each 1D profile considered separately
 - 50 degree incidence angle
 - Frequency and water content dependence
 - 1 Profile for accuracy and timing analysis
- Simulation Set 2
 - 12 separate 1D profiles
 - Lighter water content
 - Incidence angle dependence
 - 40 to 60 degrees in 5 degree steps
 - 2 cases presented for 18.7 GHz



NINDS









Simulation Set 1: Snow

NINDS



Accuracy: Error









Accuracy: Magnitude









Set 1 Results: 10.7 GHz







Set 1 Results: 18.7 GHz







Set 1 Results: 37.0 GHz







Simulation Set 2: Case 1







Set 2, Case 1 Results













Set 2, Case 2 Results







Conclusions



- V and H (Q) changes
 - Due to non-spherical drops
- No 3rd Stokes contribution
 - Off angle effects average out
 - No particle type discrimination
 - No possible retrievals
 - No contribution to surface
 - Must still account for extinction
 - Did not account for canting of rain
 - Probably small effect
- Incidence angle dependence





Current work using ARTS

- Examine radiometer calibration
- Parameterize brightness temperatures vs. cloud, water vapor and sea surface temperature
- Added polarimetric surface models
- Future work for ARTS scattering
 - Adjust code to accept canted rain drops
 - Examine degradation of surface signal



Publications



Journal Papers

- Adams, I. S., Hennon, C. C., Jones, W. L., and K. Ahmad, "Evaluation of hurricane ocean vector winds from WindSat," *IEEE Trans. Geosci. Rem. Sens.* Vol, 44, March 2006.
- Ahmad, K.A., Jones, W.L., Kasparis, T., Wiechecki Vergara, S., Adams, I.S., Park, J.-D., "Oceanic rain rate estimates from the QuikSCAT Radiometer, A Global Precipitation Mission path finder," *J. Geophys. Res.*, Vol 110, June 2005.



Publications



Conference presentations/proceedings

- Adams, I. S., Bettenhausen, M. H., Gaiser, P. W., Jones, L., "The effect of rain on WindSat pola rimetric radiometer," IGARSS '06, July 2006.
- Adams, I. S., Jones, W. L., Soisuvarn, S. Vasudevan, S. Laupattarakasem, P., "Hurricane wind retrievals using the SeaWinds scatterometer on QuikSCAT," MTS/IEEE Oceans '05, Sept. 2005.
- Adams, I. S., Jelenak, Z., Hennon, C.C., Jones W.L., "Evaluation of WindSat ocean vector wind retrievals in tropical cyclones," 27th conference on Hurricane and Tropical Meteorology, April 2006.
- Adams, I. S., Hennon, C.C. Jones, W. L., Ahmad, K., "Hurricane wind vector estimates from WindSat polarimetric radiometer," IGARSS '05, July 2005.
- Adams, I. S., Jones W. L., "High quality wind retrievals for hurricanes Isabel and Fabian using the SeaWinds scatterometer," IGARSS '04, Sept. 2004.
- Adams, I. S., Jones, W. L., Park, J. D., and T. Kasparis, "Combined active/passive hurricane retrieval algorithm for the SeaWinds Scatterometer," IGARSS '03, July 2003.
- Adams, I. S., Jones, W. L. and J. D. Park, "Improved hurricane wind speed algorithm for the SeaWinds Satellite Scatterometer," NASA Ocean Vector Wind Science Team Meeting, Jan. 14-16, 2003, Oxnard, CA.
- Adams, I. S., *et al.*, "Improved hurricane wind speed algorithm for the SeaWinds scatterometer," MTS/IEEE Oceans '02, Oct. 2002.